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# Stability of periodic travelling shallow-water waves determined by Newton's equation 

Sevdzhan Hakkaev ${ }^{1,2}$, Iliya D Iliev ${ }^{3}$ and Kiril Kirchev ${ }^{3}$<br>${ }^{1}$ Faculty of Mathematics and Informatics, Shumen University, 9712 Shumen, Bulgaria<br>${ }^{2}$ Department of Mathematics, IME-USP, Sao Paulo, Brazil<br>${ }^{3}$ Institute of Mathematics and Informatics, Bulgarian Academy of Sciences, 1113 Sofia, Bulgaria<br>E-mail: shakkaev@fmi.shu-bg.net, iliya@math.bas.bg and kpkirchev@abv.bg

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#### Abstract

We study the existence and stability of periodic travelling-wave solutions for generalized Benjamin-Bona-Mahony and Camassa-Holm equations. To prove orbital stability, we use the abstract results of Grillakis-Shatah-Strauss and the Floquet theory for periodic eigenvalue problems.


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## 1. Introduction

Consider the following equation:

$$
\begin{equation*}
u_{t}+(a(u))_{x}-u_{x x t}=\left(b^{\prime}(u) \frac{u_{x}^{2}}{2}+b(u) u_{x x}\right)_{x} \tag{1.1}
\end{equation*}
$$

where $a, b: \mathbb{R} \rightarrow \mathbb{R}$ are smooth functions and $a(0)=0$. In this paper, we study the problems of the existence and stability of periodic travelling-wave solutions $u=\varphi(x-v t)$ for (1.1). It is easy to see that whatever $a, b$ is, the equation for $\varphi$ has no dissipative terms. Hence, any travelling-wave solution of (1.1) is determined from Newton's equation which we will write below in the form $\varphi^{\prime 2}=U(\varphi)$. Therefore, by using the well-known properties of the phase portrait of Newton's equation in the $\left(\varphi, \varphi^{\prime}\right)$ plane, one can establish that under fairly broad conditions, (1.1) has at least one three-parameter family of periodic solutions $\varphi(y)=$ $\varphi\left(v, c_{1} ; \varphi_{0} ; y\right)$, where $c_{1}$ is a constant of integration and $\varphi_{0}=\min \varphi$ (see proposition 1). The parameters $v, c_{1}$ determine the phase portrait, while $\varphi_{0}$ serves to fix the periodic orbit within. Moreover, if $T=T\left(v, c_{1}, \varphi_{0}\right)$ is the minimal (sometimes called fundamental) period of $\varphi$, then $\varphi$ has exactly one local minimum and one local maximum in $[0, T)$. Therefore, $\varphi^{\prime}$ has just two zeros in each semi-open interval of length $T$. By the Floquet theory, this means that
$\varphi^{\prime}$ is either the second or the third eigenfunction of the periodic eigenvalue problem obtained from the second variation along $\varphi$ of an appropriate conservative functional $M(u)$. If the first case occurs, then one can use the abstract result of Grillakis-Shatah-Strauss [29] to prove orbital stability whenever $\ddot{\mathrm{d}}(v)=\left(\mathrm{d}^{2} / \mathrm{d} v^{2}\right) M(\varphi)$ is positive.

In the periodic case we deal with, it is not always so easy to determine the sign of $\ddot{\mathrm{d}}(v)$. To overcome this problem, we first establish a general result (see proposition 6) expressing $\ddot{d}(v)$ through some special line integrals along the energy level orbit $\{H=h\}$ of the Newtonian function $H(X, Y)=Y^{2}-U(X)$ which corresponds to $\varphi$. When $a, b$ are polynomials, these are complete Abelian integrals and one can apply methods from algebraic geometry (PicardFuchs equations, etc) to determine the possible values of $v, c_{1}$ and $\varphi_{0}$ where $\ddot{\mathrm{d}}(v)$ changes sign. Let us mention that even for $v$ and $c_{1}$ fixed, the sign of $\ddot{\mathrm{d}}(v)$ might depend on the amplitude of $\varphi$ (ruled by $\varphi_{0}$ ) as shown in proposition 8 . In this connection, we calculate explicitly the main term of $\ddot{\mathrm{d}}(v)$ in the case of arbitrary small-amplitude periodic solutions $\varphi$ of (1.1), see formula (7.7). It is shown that the main term depends on the first two isochronous constants related to the centre $\left(X_{0}, 0\right)$ into which the orbit ( $\varphi, \varphi^{\prime}$ ) shrinks when $\varepsilon=\max \varphi-\min \varphi \rightarrow 0$, and on $X_{0}$ itself as well.

We apply our results to prove orbital stability for several particular examples.
Theorem I (the modified BBM equation). Let $a(u)=2 \omega u+\beta u^{3}, b(u)=0, \beta>0$ and $u=\varphi(x-v t)$ where $v>0, \varphi(y)=\varphi\left(v, 0 ; \varphi_{0} ; y\right)$ be a periodic travelling-wave solution of (1.1) which does not oscillate around zero. Then $\varphi$ is orbitally stable in any of the cases:
(i) $3 v^{2}-8 \omega^{2} \geqslant 0$;
(ii) $3 v^{2}-8 \omega^{2}<0,2 v^{2}-2 \omega v-\omega^{2}>0$ and the period of $\varphi$ is sufficiently large.

Theorem II (the perturbed single-power BBM equation). Let $a(u)=\beta u^{2}, b(u)=\gamma \beta u, \beta>$ 0 , and let $u=\varphi(x-v t)$ where $v>0, \varphi(y)=\varphi\left(v, 0 ; \varphi_{0} ; y\right)$ be a periodic travelling-wave solution of (1.1). Then $\varphi$ is orbitally stable for small $|\gamma|$.

Theorem III (the perturbed single-power mBBM equation). Let $a(u)=\beta u^{3}, b(u)=$ $\gamma \beta u^{2}, \beta>0$, and let $u=\varphi(x-v t)$ where $v>0, \varphi(y)=\varphi\left(v, 0 ; \varphi_{0} ; y\right)$ be a periodic travelling-wave solution of (1.1) which does not oscillate around zero. Then $\varphi$ is orbitally stable for small $|\gamma|$.

Theorem IV (small-amplitude waves of the perturbed BBM equation). Let $a(u)=$ $2 \omega u+\frac{3}{2} u^{2}, b(u)=\gamma g(u)$ and $u=\varphi(x-v t)$ where $v>0, \varphi(y)=\varphi\left(v, c_{1} ; \varphi_{0} ; y\right)$ be a periodic travelling-wave solution of (1.1) having a small amplitude. Then $\varphi$ is orbitally stable for small $|\gamma|$ and $\left(\omega / v, c_{1} / v^{2}\right)$ taken in appropriate domain $\Omega \subset \mathbb{R}^{2}$.

Theorem $\mathbf{V}$ (small-amplitude waves of the perturbed mBBM equation). Let $a(u)=$ $2 \omega u+\beta u^{3}, b(u)=\gamma g(u), \beta>0$ and $u=\varphi(x-v t)$ where $v>0, \varphi(y)=\varphi\left(v, 0 ; \varphi_{0} ; y\right)$ be a periodic travelling-wave solution of (1.1) which has a small amplitude and does not oscillate around zero. Then $\varphi$ is orbitally stable for $3 v^{2}-8 \omega^{2}>0$ and small $|\gamma|$.

We point out that, unlike the other cases, in theorem IV the constant of integration $c_{1}$ is not fixed, therefore we consider the whole family of small-amplitude waves. The explicit expression of $\Omega$ is given in the proof. In theorems IV and $V, g(u)$ is a certain (fixed) function.

Let us mention that for $a(u)=2 k u+\frac{3}{2} u^{2}$ and $b(u)=u$, equation (1.1) becomes the well-known Camassa-Holm equation

$$
\begin{equation*}
u_{t}+2 k u_{x}+3 u u_{x}-u_{x x t}=2 u_{x} u_{x x}+u u_{x x x} \tag{1.2}
\end{equation*}
$$

Equation (1.2) was derived as a bi-Hamiltonian generalization of the Korteweg-de Vries equation [27] and later Camassa and Holm [11] recovered it as a water-wave model. The Camassa-Holm equation is locally well-posed in $H^{s}$ for $s>\frac{3}{2}$. Moreover, while some solutions of equation (1.2) are global, others blow up in finite time (in both the periodic and non-periodic cases) [12, 13, 16, 17, 18, 43, 49]. The solitary waves of Camassa-Holm equation are smooth in the case $k>0$ and peaked for $k=0$. Their stability is considered in [20-22, 30, 31, 41, 42].

For $a(u)=2 k u+\frac{3}{2} u^{2}$ and $b(u)=\gamma u$, equation (1.1) serves as a model equation for mechanical vibrations in a compressible elastic rod [25, 26]. Some problems such as wellposedness and blowing-up in this case were studied in [48, 49, 51], and stability of solitary waves was investigated in [23, 38].

If $b(u)=0$ in (1.1), one obtains the generalized Benjamin-Bona-Mahony (gBBM) equation for surface waves in a channel [8]. All solutions are global and their solitary waves are stable or unstable depending on $a(u)[34,46]$.

For equation (1.1), the well-posedness and stability of solitary waves in the case $a(u)=2 k u+\frac{p+2}{2} u^{p+1}$ and $b(u)=u^{p}$ are studied in [30]. For $k=0$, equation (1.1) admits peaked solitary wave solutions, which are stable (see [31]).

All these equations appear as models in the shallow-water approximation. A solitary wave is a localized steady two-dimensional gravity wave of elevation propagating on water surface. Some substantial developments in the analytical theory of water waves were discussed in the review paper by Toland [47], see also the recent papers [15, 19] describing the nature of the particle trajectories in the fluid as the solitary wave propagates on the free water surface.

The existence and stability of periodic travelling waves for nonlinear evolution equations has received a little attention. In [45] and [42], by using the integrability of the KdV and CH equations, respectively, the stability of smooth periodic solutions to these equations was proved. In [37] is shown that for $b(u)=u$, the only integrable case is the CH equation. Therefore, one cannot use the approach proposed by McKean [45] to handle the much more general case we deal with.

Recently, Angulo, Bona and Scialom [4] developed a complete theory of the stability of cnoidal waves for the KdV equation. The solution $u(x, t)=\varphi_{c}(x-c t)$ of KdV satisfies the equation

$$
\varphi_{c}^{\prime \prime}+\frac{1}{2} \varphi_{c}^{2}-c \varphi_{c}=A_{\varphi_{c}},
$$

where $A_{\varphi_{c}}$ is an integration constant. An explicit form for $\varphi_{c}$ in the periodic case is

$$
\varphi_{c}(\xi)=\beta_{2}+\left(\beta_{3}-\beta_{2}\right) c n^{2}\left(\sqrt{\frac{\beta_{3}-\beta_{1}}{12}} \xi ; k\right)
$$

where $c n$ is the Jacobi elliptic function and the following relations take place:
$\beta_{1}<\beta_{2}<\beta_{3}, \quad k^{2}=\frac{\beta_{3}-\beta_{2}}{\beta_{3}-\beta_{1}}, \quad \beta_{1}+\beta_{2}+\beta_{3}=3 c, \quad A_{\varphi_{c}}=-\frac{1}{6} \Sigma_{i<j} \beta_{i} \beta_{j}$.
Since $c n(u+2 K)=-c n(u)$ where $K=K(k)=\int_{0}^{1} \frac{\mathrm{~d} t}{\sqrt{\left(1-t^{2}\right)\left(1-k^{2} t^{2}\right)}}$ is the complete elliptic integral of the first kind, then $\varphi_{c}$ has the fundamental (i.e. minimal) period $T_{\varphi_{c}}$ given by

$$
T_{\varphi_{c}}=\frac{4 \sqrt{3}}{\sqrt{\beta_{3}-\beta_{1}}} K(k)
$$

Moreover, $T_{\varphi_{c}}$ depends on the speed $c$ and satisfies the inequality

$$
T_{\varphi_{c}}^{2}>\frac{(2 \pi)^{2}}{\sqrt{c^{2}+2 A_{\varphi_{c}}}}
$$

Using the abstract results of Grillakis, Shatah and Strauss (adapted to the periodic context), the authors proved in [4] that the cnoidal waves having a zero mean-value are orbitally stable.

Other new explicit formulae for the periodic travelling waves based on the Jacobi elliptic function of type dnoidal, together with their stability, have been obtained by Angulo [2, 3] for the nonlinear Schrödinger (NLS) equation $\mathrm{i} u_{t}+u_{x x}+|u|^{2} u=0$, modified KdV equation $u_{t}+3 u^{2} u_{x}+u_{x x x}=0$ and Hirota-Satsuma system

$$
\left\{\begin{array}{l}
u_{t}-a\left(u_{x x x}+6 u u_{x}\right)=2 b v v_{x} \\
v_{t}+v_{x x x}+3 u v_{x}=0 .
\end{array}\right.
$$

For NLS the solutions are of the form $u(x, t)=\mathrm{e}^{\mathrm{i} w t} \varphi_{w}(x)$, where $\varphi_{w}(x)$ is a periodic smooth function with period $L>0$. The solution $\varphi_{w}$ is of the form

$$
\varphi_{w}(x)=\eta_{1} \mathrm{~d} n\left(\frac{\eta_{1}}{\sqrt{c}} x ; k\right)
$$

where $\eta_{1}$ and the modulus $k$ depend smoothly on $w$. The orbit $\Omega_{\varphi_{w}}$,

$$
\Omega_{\varphi_{w}}=\left\{\mathrm{e}^{\mathrm{i} \theta} \varphi_{w}(\cdot+y),(y, \theta) \in \mathbb{R} \times[0,2 \pi)\right\}
$$

generated by the dnoidal wave $\varphi_{w}$ is stable by perturbation of periodic function with period $L$ and nonlinearly unstable by perturbation of periodic function with period $2 L$. In all these works, it was necessary to use an elaborated spectral theory for the periodic eigenvalue problem,

$$
\left\{\begin{array}{l}
\frac{\mathrm{d}^{2}}{\mathrm{~d} x^{2}} \Psi+\left[\rho-n(n+1) k^{2}{\left.s n^{2}(x ; k)\right] \Psi=0}^{\Psi(0)=\Psi(2 K(k)), \quad \Psi^{\prime}(0)=\Psi^{\prime}(2 K(k))}\right. \tag{1.3}
\end{array}\right.
$$

with specific values of $n \in \mathbb{N}$. We will also make use of systems similar to (1.3).
The present paper is organized as follows. In section 2 , we formulate and sketch the proof of a local well-posedness result for equation (1.1) in periodic $H^{s}$ spaces. In section 3, we prove the existence of periodic travelling waves of a given (admissible) period and study their properties. In section 4, we prove the orbital stability result under some hypotheses (see assumption 1). In section 5, several particular examples are considered. For most of them, we determine $\varphi$ explicitly and show that $\varphi^{\prime}$ is the second eigenfunction of the respective periodic eigenvalue problem. We also determine the sign of $\ddot{\mathrm{d}}(v)$ to outline the cases satisfying assumption 1 . In section 6, perturbation theory is applied to prove the orbital stability of periodic travelling waves for generalized Benjamin-Bona-Mahony and generalized Camassa-Holm equations (the case of small $b$ in the right-hand side of (1.1)). In section 7, we study the small-amplitude periodic travelling waves of (1.1) and determine the sign of $\ddot{\mathrm{d}}(v)$ for them.

## 2. Local well-posedness

In this section, we discuss the local well-posedness of the Cauchy problem for equation (1.1). We begin by introducing some notation and by recalling related definitions we shall use throughout the paper.

Let $\mathcal{P}=C_{\text {per }}^{\infty}$ denote the collection of all functions which are $C^{\infty}$ and periodic with a period $T>0$. The topological dual of $\mathcal{P}$ will be denoted by $\mathcal{P}^{\prime}$. If $\Psi \in \mathcal{P}^{\prime}$ then we denote by $\Psi(f)=\langle\Psi, f\rangle$ the value of $\Psi$ at $f$. Define the functions $\Theta_{k}(x)=\exp (2 \pi \mathrm{i} k x / T), k \in \mathbb{Z}$. The Fourier transform of $\Psi \in \mathcal{P}^{\prime}$ is the function $\widehat{\Psi}: \mathbb{Z} \rightarrow \mathbb{C}$ defined by $\widehat{\Psi}(k)=\frac{1}{T}\left\langle\Psi, \Theta_{-k}\right\rangle$. If $\Psi$ is a periodic function with a period $T$, we have

$$
\widehat{\Psi}(k)=\frac{1}{T} \int_{0}^{T} \Psi(x) \exp (-2 \pi \mathrm{i} k x / T) \mathrm{d} x
$$

For $s \in \mathbb{R}$, the Sobolev space $H^{s}([0, T])$ is the set of all $f \in \mathcal{P}^{\prime}$ such that

$$
\|f\|_{s}^{2}=T \sum_{k=-\infty}^{\infty}\left(1+|k|^{2}\right)^{s}|\widehat{f}(k)|^{2}<\infty
$$

Certainly, $H^{s}([0, T])$ is a Hilbert space with respect to the inner product

$$
(f, g)_{s}=T \sum_{k=-\infty}^{\infty}\left(1+|k|^{2}\right)^{s} \widehat{f}(k) \overline{\widehat{g}(k)}
$$

Since $H^{s}([0, T]) \subset L^{2}([0, T])$ for every $s \geqslant 0$, we obtain via Plancherel identity that for every $n \in \mathbb{N}$

$$
\|f\|_{n}^{2}=\sum_{j=0}^{n}\left\|f^{(j)}\right\|^{2}
$$

where $f^{(j)}$ represents the $j$ th derivative of $f$ taken in the sense of $\mathcal{P}^{\prime}$. Moreover, Sobolev's lemma states that if $s>l+\frac{1}{2}$, then $H^{s}([0, T]) \hookrightarrow C_{\text {per }}^{l}$ where

$$
C_{\mathrm{per}}^{l}=\left\{f \in C^{l}: f^{(j)} \text { is periodic with a period } T \text { for } j=0, \ldots, l\right\} .
$$

One can prove the following result about equation (1.1).
Theorem 1. Assume that $a, b \in C^{m+3}(\mathbb{R}), m \geqslant 2$. Given $u_{0} \in H^{s}, \frac{3}{2}<s<m$, there exists $a$ maximal $t_{0}>0$ and a unique solution $u(x, t)$ to (1.1) such that

$$
u \in C\left(\left[0, t_{0}\right), H^{s}\right) \cap C^{1}\left(\left[0, t_{0}\right), H^{s-1}\right)
$$

Moreover, the solution depends continuously on the initial data.
Proof. Take $u \in H^{s}$ and let

$$
A(u)=b(u) \partial_{x}, f(u)=\left(1-\partial_{x}^{2}\right)^{-1}\left[b(u) u_{x}-\partial_{x}\left(\frac{1}{2} b^{\prime}(u) u_{x}^{2}+a(u)\right)\right] .
$$

Using the above notations, one can rewrite equation (1.1) in the following form:

$$
u_{t}+A(u) u=f(u)
$$

In a similar way as in theorem 2.2 in [31] (dealing with the non-periodic case), we have
(1) $A(u)$ is quasi-m-accretive, uniformly on the bounded sets in $H^{s-1}$. Moreover, $A(u) \in$ $L\left(H^{s}, H^{s-1}\right)$ (where $L(X, Y)$ is the space of all linear bounded operators from $X$ to $Y, L(X)=L(X, X))$ and

$$
\left\|\left(A\left(u_{1}\right)-A\left(u_{2}\right)\right) u_{3}\right\|_{s-1} \leqslant \mu_{1}\left\|u_{1}-u_{2}\right\|_{s-1}\left\|u_{3}\right\|_{s} .
$$

(2) Define $\Lambda=\left(1-\partial_{x}^{2}\right)^{1 / 2}, B(u)=\left[\Lambda, b(u) \partial_{x}\right] \Lambda^{-1}$ for $u \in H^{s}$, where $[\Lambda, M]$ denotes the commutator of $\Lambda$ and M . Then $B(u) \in L\left(H^{s-1}\right)$ and
$\left\|\left(B\left(u_{1}\right)-B\left(u_{2}\right)\right) u_{3}\right\|_{s-1} \leqslant \mu_{2}\left\|u_{1}-u_{2}\right\|_{s}\left\|u_{3}\right\|_{s-1}, \quad u_{1}, u_{2} \in H^{s}, \quad u_{3} \in H^{s-1}$.
(3) $f(u)$ is bounded on the bounded sets in $H^{s}$ and satisfies

$$
\begin{array}{ll}
\left\|f\left(u_{1}\right)-f\left(u_{2}\right)\right\|_{s} \leqslant \mu_{3}\left\|u_{1}-u_{2}\right\|_{s}, & u_{1}, u_{2} \in H^{s} \\
\left\|f\left(u_{1}\right)-f\left(u_{2}\right)\right\|_{s-1} \leqslant \mu_{4}\left\|u_{1}-u_{2}\right\|_{s-1}, & u_{1}, u_{2} \in H^{s-1}
\end{array}
$$

Applying Kato's theory for abstract quasilinear evolution equations [40], we obtain the local well-posedness of equation (1.1) in $H^{s}$, for $\frac{3}{2}<s<m$. The solution $u(x, t)$ belongs to $C\left(\left[0, t_{0}\right), H^{s}\right) \cap C^{1}\left(\left[0, t_{0}\right), H^{s-1}\right)$.

It should be noticed that a local well-posedness result is the best one can hope for. Indeed, in the BBM case all solutions are global in time, but in the Camassa-Holm case some solutions are global and some develop singularities in finite time, depending on the initial data. The singularities are of wave-breaking type, that is the solution remains bounded but its slope becomes unbounded in finite time [13, 14, 16-18]. In the non-periodic case, a continuation of the solutions of the CH equation after wave breaking was established by Bressan and Constantin [10].

## 3. Periodic travelling-wave solutions

We are looking for a travelling-wave solution of (1.1) of the form $u(x, t)=\varphi(x-v t)$. We assume that $\varphi$ is smooth and bounded in $\mathbb{R}$. The following two cases appear:
(i) $\varphi^{\prime} \neq 0$ in $\mathbb{R}$ and $\varphi_{-}<\varphi<\varphi_{+}$(corresponding to the kink-wave solution);
(ii) $\varphi^{\prime}(\xi)=0$ for some $\xi \in \mathbb{R}$. Denote $\varphi_{0}=\varphi(\xi), \varphi_{2}=\varphi^{\prime \prime}(\xi)$.

Below we will deal with the second case. Replacing in (1.1) we get

$$
\begin{equation*}
-v \varphi^{\prime}+(a(\varphi))^{\prime}+v \varphi^{\prime \prime \prime}=\left(b^{\prime}(\varphi) \frac{\varphi^{\prime 2}}{2}+b(\varphi) \varphi^{\prime \prime}\right)^{\prime} \tag{3.1}
\end{equation*}
$$

By integrating (3.1) twice, one obtains

$$
\begin{align*}
& -v \varphi+a(\varphi)+v \varphi^{\prime \prime}=b^{\prime}(\varphi) \frac{\varphi^{\prime 2}}{2}+b(\varphi) \varphi^{\prime \prime}+c_{1}  \tag{3.2}\\
& {[v-b(\varphi)] \frac{\varphi^{\prime 2}}{2}=c_{2}+c_{1} \varphi+\frac{v}{2} \varphi^{2}-A(\varphi), \quad A(\varphi)=\int_{0}^{\varphi} a(s) \mathrm{d} s} \tag{3.3}
\end{align*}
$$

with some constants $c_{1}, c_{2}$. In case (ii), one has respectively

$$
\begin{aligned}
& c_{1}=a\left(\varphi_{0}\right)-v \varphi_{0}+\left(v-b\left(\varphi_{0}\right)\right) \varphi_{2} \\
& c_{2}=A\left(\varphi_{0}\right)+\frac{1}{2} v \varphi_{0}^{2}-\varphi_{0} a\left(\varphi_{0}\right)-\left(v-b\left(\varphi_{0}\right)\right) \varphi_{0} \varphi_{2}=A\left(\varphi_{0}\right)-\frac{1}{2} v \varphi_{0}^{2}-c_{1} \varphi_{0}
\end{aligned}
$$

Next, we are going to look for periodic travelling-wave solutions $\varphi$. Consider in the plane $(X, Y)=\left(\varphi, \varphi^{\prime}\right)$ the Hamiltonian system

$$
\begin{align*}
& \dot{X}=[v-b(X)] Y=H_{Y}, \\
& \dot{Y}=\frac{1}{2} b^{\prime}(X) Y^{2}-a(X)+v X+c_{1}=-H_{X}, \tag{3.4}
\end{align*}
$$

with a Hamiltonian function

$$
H(X, Y)=[v-b(X)] \frac{Y^{2}}{2}+A(X)-\frac{v}{2} X^{2}-c_{1} X
$$

Then (3.3) becomes $H\left(\varphi, \varphi^{\prime}\right)=c_{2}$ and the curve $s \rightarrow\left(\varphi\left(s-s_{0}\right), \varphi^{\prime}\left(s-s_{0}\right)\right)$ determined by (3.3) lies on the energy level $H=c_{2}$ of the Hamiltonian $H(X, Y)$. Within the analytical class, system (3.4) has periodic solutions if and only if it has a centre. Each centre is surrounded by a continuous band of periodic trajectories (called period annulus) which terminates at a certain separatrix contour on the Poincaré sphere. As far as the straight lines $X=X_{0}$ where $v-b\left(X_{0}\right)=0$ are (unions of) trajectories, the critical points of centre type of (3.4) are given by the critical points on $Y=0$ having a negative Hessian. These are the points $\left(X_{0}, 0\right)$, where

$$
\begin{equation*}
c_{1}+v X_{0}-a\left(X_{0}\right)=0, \quad\left[v-b\left(X_{0}\right)\right]\left[v-a^{\prime}\left(X_{0}\right)\right]<0 . \tag{3.5}
\end{equation*}
$$

(For simplicity, we will not consider here the case of a degenerate centre when the Hessian becomes zero.)

The above considerations lead us to the following statement.
Proposition 1. Let $c_{1}$ and $v$ be constants such that conditions (3.5) are satisfied for some $X_{0} \in \mathbb{R}$. Then there is an open interval $\Delta$ containing $X_{0}$, such that
(i) for any $\varphi_{0} \in \Delta, \varphi_{0}<X_{0}$, the solution of (3.1) satisfying

$$
\varphi(\xi)=\varphi_{0}, \quad \varphi^{\prime}(\xi)=0, \quad \varphi^{\prime \prime}(\xi)=\frac{c_{1}+v \varphi_{0}-a\left(\varphi_{0}\right)}{v-b\left(\varphi_{0}\right)}
$$

is periodic;
(ii) if $\varphi_{1} \in \Delta, \varphi_{1}>X_{0}$ is the nearest to $X_{0}$ solution of $H(X, 0)=H\left(\varphi_{0}, 0\right)$, then $\varphi_{0} \leqslant \varphi \leqslant \varphi_{1}$;
(iii) if $T$ is the minimal period of $\varphi$, then in each interval $[s, s+T)$ the function $\varphi$ has just one minimum and one maximum (equal to $\varphi_{0}$ and $\varphi_{1}$, respectively) and is strictly monotone elsewhere.

Proof. Part (i) follows from the analysis already done above. Parts (ii) and (iii) follow from the fact that $H(X, Y)$ is symmetric with respect to $Y$ and because for any fixed $\bar{X}$ the equation $H(\bar{X}, Y)=$ const has just two solutions $\bar{Y},-\bar{Y}$.

Example. Assume that $a, b$ are polynomials and $\operatorname{deg} a=2, \operatorname{deg} b \leqslant 1$. For $c_{1}, v$ properly chosen, the quadratic equation in (3.5) will have two distinct real roots $X_{1}<X_{2}$. It is easy to see that if $v-b(X) \neq 0$ in [ $X_{1}, X_{2}$ ], then (3.5) holds for just one of these roots. If $a^{\prime \prime} b^{\prime}<0$ and $v-b(X)$ vanishes at $X_{1}$, then (3.5) holds for $X_{2}$, and vice versa. Finally, if $a^{\prime \prime} b^{\prime}<0$ and $v-b(X)$ has a root within $\left(X_{1}, X_{2}\right)$, then both $X_{1}$ and $X_{2}$ satisfy (3.5). To summarize, the phase portraits of all quadratic systems (3.4) (including the exceptional case $\operatorname{deg} a \leqslant 1$, $\operatorname{deg} b=1$ ) satisfying (3.5), that is having a centre, are divided into seven topologically different classes, see, e.g. [32].

Remark 1. Below, we shall denote $\Delta^{-}=\Delta \cap\left\{\left(-\infty, X_{0}\right)\right\}, \Delta^{+}=\Delta \cap\left\{\left(X_{0}, \infty\right)\right\}$. It is possible that (3.3) also has periodic solutions for initial values $\varphi_{0}$ far from $X_{0}$. To study them, it is needed to specify the functions $a$ and $b$ in (1.1).

Let us denote
$U(s)=\frac{2 c_{2}+2 c_{1} s+v s^{2}-2 A(s)}{v-b(s)}=\frac{2 A\left(\varphi_{0}\right)-v \varphi_{0}^{2}-2 c_{1} \varphi_{0}+2 c_{1} s+v s^{2}-2 A(s)}{v-b(s)}$.
Then for $\varphi_{0} \leqslant \varphi \leqslant \varphi_{1}$ one can rewrite (3.3) as $\varphi^{\prime}(\sigma)=\sqrt{U(\varphi(\sigma))}$. Integrating the equation along the interval $[\xi, s] \subset[\xi, \xi+T / 2]$ yields an implicit formula for the value of $\varphi(s)$ :

$$
\begin{equation*}
\int_{\varphi_{0}}^{\varphi(s)} \frac{\mathrm{d} \sigma}{\sqrt{U(\sigma)}}=s-\xi, \quad s \in[\xi, \xi+T / 2] \tag{3.6}
\end{equation*}
$$

For $s \in[\xi+T / 2, \xi+T]$ one has $\varphi(s)=\varphi(T+2 \xi-s)$. We recall that the period function $T$ of a Hamiltonian flow generated by $H_{0} \equiv \frac{1}{2} Y^{2}-\frac{1}{2} U(X)=0$ is determined from

$$
\begin{equation*}
T=\int_{0}^{T} \mathrm{~d} t=\oint_{H_{0}=0} \frac{\mathrm{~d} X}{Y}=2 \int_{\varphi_{0}}^{\varphi_{1}} \frac{\mathrm{~d} X}{\sqrt{U(X)}} \tag{3.7}
\end{equation*}
$$

This is in fact the derivative (with respect to the energy level) of the area surrounded by the periodic trajectory through the point $\left(\varphi_{0}, 0\right)$ in the $(X, Y)=\left(\varphi, \varphi^{\prime}\right)$ plane.

Consider the continuous family of periodical travelling-wave solutions $\{u=\varphi(x-v t)\}$ of (1.1) and (3.3) going through the points $\left(\varphi, \varphi^{\prime}\right)=\left(\varphi_{0}, 0\right)$ where $\varphi_{0} \in \Delta^{-}$. For any $\varphi_{0} \in \Delta^{-}$, denote by $T=T\left(\varphi_{0}\right)$ the corresponding period. One can see (e.g., by using formula (3.7)) that the period function $\varphi_{0} \rightarrow T\left(\varphi_{0}\right)$ is smooth. To check this, it suffices to perform a change of the variable

$$
\begin{equation*}
X=\frac{\varphi_{1}-\varphi_{0}}{2} s+\frac{\varphi_{1}+\varphi_{0}}{2} \tag{3.8}
\end{equation*}
$$

in integral (3.7) and use that

$$
\begin{equation*}
U\left(\varphi_{0}\right)=U\left(\varphi_{1}\right)=0 \tag{3.9}
\end{equation*}
$$

Also, it is not difficult to verify (see section 7) that

$$
T\left(\varphi_{0}\right) \rightarrow T_{0}=2 \pi \sqrt{\frac{v-b\left(X_{0}\right)}{a^{\prime}\left(X_{0}\right)-v}} \quad \text { as } \quad \varphi_{0} \uparrow X_{0}
$$

Conversely, taking $v, c_{1}$ to satisfy the conditions of proposition 1 and fixing $T$ in a proper interval, one can determine $\varphi_{0}$ and $\varphi_{1}$ as smooth functions of $v, c_{1}$ so that the periodic solution $\varphi$ given by (3.6) will have a period $T$. The condition for this is the monotonicity of the period.

Definition 1. We say that the period $T=T\left(\varphi_{0}\right)$ is not critical provided that $T^{\prime}\left(\varphi_{0}\right) \neq 0$.
If the period $T\left(\varphi_{0}\right)$ is not critical for any $\varphi_{0} \in \Delta_{-}$, then the period function is strictly monotone along the period annulus and its range is an open interval $I$ having $T_{0}$ as an endpoint.

Proposition 2. Let $\mathcal{A}$ be a period annulus of (3.4) which surrounds a nondegenerate centre and has a monotone period function. Then for any $T \in I$ there is a unique $\varphi_{0} \in \Delta_{-}$satisfying $T\left(\varphi_{0}\right)=T$. Its derivative $\dot{\varphi}_{0}$ with respect to $v$ is determined from

$$
\begin{equation*}
\dot{\varphi}_{0}\left[c_{1}+v \varphi_{0}-a\left(\varphi_{0}\right)\right] \frac{\mathrm{d}}{\mathrm{~d} h} \oint_{H=h} \frac{\mathrm{~d} x}{y}=\frac{\mathrm{d}}{\mathrm{~d} h} \oint_{H=h} \frac{\left(x^{2}+y^{2}-\varphi_{0}^{2}\right) \mathrm{d} x}{2 y} \tag{3.10}
\end{equation*}
$$

where

$$
\begin{equation*}
h=A\left(\varphi_{0}\right)-\frac{v}{2} \varphi_{0}^{2}-c_{1} \varphi_{0} . \tag{3.11}
\end{equation*}
$$

Proof. We use the implicit function theorem (IFT) and the Gelfand-Leray form (see [6, chapter 3]). Let $I$ be the range of the period function along $\mathcal{A}$. For $T \in I$ and $\varphi_{0} \in \Delta_{-}$, denote

$$
G\left(v, c_{1}, \varphi_{0}\right)=\bar{G}\left(v, c_{1}, h\right)=T\left(\varphi_{0}\right)-T=\oint_{H=h} \frac{\mathrm{~d} x}{y}-T
$$

Then

$$
0 \neq T^{\prime}\left(\varphi_{0}\right)=\frac{\partial G}{\partial \varphi_{0}}=\frac{\mathrm{d} h}{\mathrm{~d} \varphi_{0}} \frac{\partial \bar{G}}{\partial h}=\left[a\left(\varphi_{0}\right)-v \varphi_{0}-c_{1}\right] \frac{\mathrm{d}}{\mathrm{~d} h} \oint_{H=h} \frac{\mathrm{~d} x}{y}
$$

and the IFT works. Hence

$$
\begin{equation*}
\bar{G}=\frac{\partial \bar{G}}{\partial h} \dot{h}+\frac{\partial \bar{G}}{\partial v}=0 \tag{3.12}
\end{equation*}
$$

From $H(x, y)=h$, we obtain the covariant derivatives

$$
\begin{equation*}
(v-b(x)) y \frac{\mathrm{~d} y}{\mathrm{~d} h}=1, \quad \frac{1}{2}\left(y^{2}-x^{2}\right)+(v-b(x)) y \dot{y}=0 \tag{3.13}
\end{equation*}
$$

Therefore, by using the Gelfand-Leray form to calculate the derivatives, we get
$\bar{G}\left(v, c_{1}, h\right)=\frac{\mathrm{d}}{\mathrm{d} h} \oint_{H=h}(v-b(x)) y \mathrm{~d} x-T, \quad \frac{\partial \bar{G}}{\partial h}=\frac{\mathrm{d}}{\mathrm{d} h} \oint_{H=h} \frac{\mathrm{~d} x}{y}$,

$$
\begin{equation*}
\frac{\partial \bar{G}}{\partial v}=\frac{\mathrm{d}}{\mathrm{~d} h} \oint_{H=h}[y+(v-b(x)) \dot{y}] \mathrm{d} x=\frac{\mathrm{d}}{\mathrm{~d} h} \oint_{H=h} \frac{\left(x^{2}+y^{2}\right) \mathrm{d} x}{2 y} \tag{3.15}
\end{equation*}
$$

Hence, by (3.12), (3.14) and (3.15), we obtain

$$
\begin{equation*}
\dot{h} \frac{\mathrm{~d}}{\mathrm{~d} h} \oint_{H=h} \frac{\mathrm{~d} x}{y}+\frac{\mathrm{d}}{\mathrm{~d} h} \oint_{H=h} \frac{\left(x^{2}+y^{2}\right) \mathrm{d} x}{2 y}=0 \tag{3.16}
\end{equation*}
$$

Finally, from (3.11) one obtains $\dot{h}=-\frac{1}{2} \varphi_{0}^{2}+\left[a\left(\varphi_{0}\right)-v \varphi_{0}-c_{1}\right] \dot{\varphi}_{0}$. Together with (3.16), this implies (3.10).

Remark 2. Obviously, one can formulate a local variant of proposition 2 concerning a given noncritical period $T\left(\varphi_{0}\right)$ only. As far as $T^{\prime}\left(\varphi_{0}\right) \neq 0$, the same proof clearly goes and no restrictions concerning the period annulus $\mathcal{A}$ are needed.

The perturbation result we establish below will be needed in section 6. Instead of (1.1), consider now a small perturbation of the generalized BBM equation

$$
\begin{equation*}
u_{t}+(a(u))_{x}-u_{x x t}=\gamma\left(b^{\prime}(u) \frac{u_{x}^{2}}{2}+b(u) u_{x x}\right)_{x}, \quad|\gamma| \ll 1 \tag{3.17}
\end{equation*}
$$

and let $\left\{\varphi^{\gamma}(x-v t)\right\}$ be the family of corresponding periodic travelling-wave solutions going through points $\left(\varphi_{0}^{\gamma}, 0\right)$ in the $\left(\varphi, \varphi^{\prime}\right)$ plane. Denote the related periods by $T\left(\varphi_{0}^{\gamma}\right)$.

Proposition 3. Assume that the period $T=T\left(\varphi_{0}^{0}\right)$ related to the solution $\varphi^{0}(x-v t)$ of (3.17) (with $\gamma=0$ ) is not critical. Then
(i) there is a smooth function $\gamma \rightarrow \varphi_{0}(\gamma)$ defined for small $|\gamma|$ and satisfying $\varphi_{0}(0)=\varphi_{0}^{0}$, such that the travelling-wave solution $\varphi^{\gamma}(x-v t)$ of (3.17) going through the point $\left(\varphi_{0}(\gamma), 0\right)$ is periodical and has a (minimal) period $T$;
(ii) $\max _{[0, T]}\left|\varphi^{\gamma}-\varphi^{0}\right|=O(\gamma)$.

## Proof.

(i) The proof relies on the implicit function theorem. Consider system (3.4) with $b$ replaced by $\gamma b$. Since for $\varphi_{0} \in \Delta^{-}$the point $\left(\varphi_{0}, 0\right)$ is not critical for (3.4), IFT yields that there is a smooth function $\varphi_{1}=\varphi_{1}\left(\varphi_{0}, \gamma\right)$ determined from $H\left(\varphi_{1}, 0\right)=H\left(\varphi_{0}, 0\right)$, which takes values in $\Delta^{+}$and $\varphi_{0} \leqslant \varphi^{\gamma} \leqslant \varphi_{1}$. Introducing a new variable (3.8) in (3.7) and making use of (3.9), we can rewrite (3.7) in the form $G\left(\gamma, \varphi_{0}\right)=0$ where

$$
G\left(\gamma, \varphi_{0}\right)=2 \int_{-1}^{1} \frac{\mathrm{~d} s}{\sqrt{\left(1-s^{2}\right) U_{1}\left(\gamma, \varphi_{0}, s\right)}}-T
$$

with $U_{1}$ a smooth positive function. We further have

$$
G\left(0, \varphi_{0}^{0}\right)=0,\left.\quad \frac{\mathrm{~d} G\left(\gamma, \varphi_{0}\right)}{\mathrm{d} \varphi_{0}}\right|_{\gamma=0}=T^{\prime}\left(\varphi_{0}^{0}\right) \neq 0
$$

therefore (i) follows, again by the IFT.
(ii) Consider system (3.4) with $b$ replaced by $\gamma b$ and initial data $X(\xi)=\varphi_{0}(\gamma), Y(\xi)=0$. By the uniqueness and the smooth dependence theorems, the solution $\varphi^{\gamma}$ is smooth and therefore a uniformly Lipschitz continuous function with respect to $\gamma$ for $|\gamma|$ small, on each interval $[\xi, \xi+T]$.

In support to the hypothesis of proposition 3, we state the following (in fact, known) result about the period $T=T\left(\varphi_{0}\right)$.

Proposition 4. (i) There is no critical period provided that a (u) in (3.17) is a polynomial of degree 2. (ii) There is at most one critical period, if $a(u)$ is a polynomial of degree 3.

Proof. Consider the period function related to the Hamiltonian system (3.4) (with $b(X)=0$ ) as a function $T(h)$ of the energy level $H=h$. It is well known that $T^{\prime}(h) \neq 0$ if $\operatorname{deg} a=2$ [24] and that $T^{\prime}(h)$ has at most one zero if $\operatorname{deg} a=3$ [28]. As

$$
\frac{\mathrm{d}}{\mathrm{~d} \varphi_{0}}=\left[a\left(\varphi_{0}\right)-v \varphi_{0}-c_{1}\right] \frac{\mathrm{d}}{\mathrm{~d} h}
$$

and the coefficient is not zero for $\varphi_{0} \in \Delta^{-}$according to (3.5), the claim follows.

## 4. Conservation laws and conditional stability

We now turn to our main stability problem. Take $v \in \mathbb{R}$ and denote by $\varphi_{v}$ the periodic travelling-wave solution $u=\varphi(x-v t)$ of (1.1).

Definition 2. The periodic travelling-wave solution $\varphi_{v}$ of (1.1) is said to be stable, if for every $\varepsilon>0$ there is $\delta>0$ such that if $u \in C\left(\left[0, t_{0}\right), H^{1}\right)$ is a solution of (1.1) satisfying $\left\|u(0)-\varphi_{v}\right\|_{1}<\delta$, then

$$
\inf _{r \in \mathbb{R}}\|u(\cdot, t)-\varphi(\cdot-r)\|_{1}<\varepsilon \quad \text { for every } t \in\left[0, t_{0}\right) .
$$

As usual, an important role in our construction will be played by some functionals which are invariant with respect to time $t$. Equation (1.1) possesses the following conservation laws:

$$
\begin{aligned}
& E(u)=-\int_{0}^{T}\left[A(u)+\frac{b(u)}{2} u_{x}^{2}\right] \mathrm{d} x \\
& Q(u)=\frac{1}{2} \int_{0}^{T}\left(u^{2}+u_{x}^{2}\right) \mathrm{d} x \\
& V(u)=\int_{0}^{T} u \mathrm{~d} x
\end{aligned}
$$

where $T$ is the minimal period of the solution $u(x, t)$ and $A^{\prime}(u)=a(u)$. Let us denote for short $M=E+v Q$. In terms of $E, Q$ and $V$, equation (3.2) with $\varphi=\varphi_{v}$ reads

$$
\begin{equation*}
M^{\prime}\left(\varphi_{v}\right)+c_{1}=E^{\prime}\left(\varphi_{v}\right)+v Q^{\prime}\left(\varphi_{v}\right)+c_{1} V^{\prime}\left(\varphi_{v}\right)=0 . \tag{4.1}
\end{equation*}
$$

Let

$$
\begin{equation*}
\mathrm{d}(v)=M\left(\varphi_{v}\right) \tag{4.2}
\end{equation*}
$$

Then differentiating (4.2) with respect to $v$, we obtain

$$
\begin{align*}
\dot{\mathrm{d}}(v) & =Q\left(\varphi_{v}\right), \\
\ddot{\mathrm{d}}(v) & =\left\langle Q^{\prime}\left(\varphi_{v}\right), \dot{\varphi}_{v}\right\rangle=\frac{\mathrm{d}}{\mathrm{~d} v}\left(\frac{1}{2} \int_{0}^{T}\left(\varphi_{v}^{2}+\varphi_{v}^{\prime 2}\right) \mathrm{d} x\right) . \tag{4.3}
\end{align*}
$$

Consider in $L^{2}[0, T]$ the operator $\mathcal{H}_{v}$ defined by the formal differential expression
$\mathcal{H}_{v}=M^{\prime \prime}\left(\varphi_{v}\right)=\left(b\left(\varphi_{v}\right)-v\right) \partial_{x}^{2}+b^{\prime}\left(\varphi_{v}\right) \varphi_{v}^{\prime} \partial_{x}+v-a^{\prime}\left(\varphi_{v}\right)+\frac{1}{2} b^{\prime \prime}\left(\varphi_{v}\right) \varphi_{v}^{\prime 2}+b^{\prime}\left(\varphi_{v}\right) \varphi_{v}^{\prime \prime}$.
As $\mathcal{H}_{v} \varphi_{v}^{\prime}=0$, zero is in the spectrum of $\mathcal{H}_{v}$. We make the following assumption concerning $\mathcal{H}_{v}$ and $\ddot{\mathrm{d}}(v)$.

## Assumption 1.

(i) The operator $\mathcal{H}_{v}$ has a unique negative eigenvalue, a simple eigenvalue 0 and the rest of its spectrum is positive.
(ii) $\ddot{\mathrm{d}}(v)>0$.

For $\varepsilon>0$, consider a neighbourhood $\mathcal{U}_{\varepsilon}$ in the set of all translations of $\varphi_{v}$ defined by

$$
\mathcal{U}_{\varepsilon}=\left\{u \in H^{1}([0, T]): \inf _{r \in \mathbb{R}}\left\|u-\varphi_{v}(\cdot-r)\right\|_{1}<\varepsilon\right\} .
$$

By using (4.1) and assumption 1(i), one can establish the following statement, whose proof is the same as in [29] and for this reason we omit it.

Lemma 1. There exists $\varepsilon>0$ and a $C^{1}$-map $\alpha: \mathcal{U}_{\varepsilon} \rightarrow \mathbb{R} / T$ such that for all $u \in \mathcal{U}_{\varepsilon}$ and all $r \in \mathbb{R}$, the following holds:
(i) $\|u(\cdot+\alpha(u))-\varphi\|_{1} \leqslant\|u(\cdot+r)-\varphi\|_{1}$,
(ii) $\left\langle u(\cdot+\alpha(u)), \varphi_{v}^{\prime}\right\rangle=0$,
(iii) $\quad \alpha(u(\cdot+r))=\alpha(u)-r$, modulo the period.

Proof. [See lemma 3.2 in [29]].
Note that $\alpha(u)$ is constructed to be a minimizer of the norm $\|u(\cdot+\alpha(u))-\varphi\|_{1}$. Moreover, $|\alpha(u)|$ is smaller than the half-period. For other details, see the proof of lemma 3.2 in [29].

Lemma 2. If assumption 1 holds and y satisfies $\left\langle Q^{\prime}\left(\varphi_{v}\right), y\right\rangle=\left\langle\varphi_{v}^{\prime}, y\right\rangle=0$, then there exists $K>0$ so that

$$
\left\langle\mathcal{H}_{v} y, y\right\rangle>K\|y\|_{1}^{2} .
$$

Proof. Differentiating (4.1) with respect to $v$ yields (since $c_{1}$ does not depend on $v$ )

$$
\begin{equation*}
\mathcal{H}_{v} \dot{\varphi}_{v}=-Q^{\prime}\left(\varphi_{v}\right) \tag{4.5}
\end{equation*}
$$

and from (4.3) we obtain

$$
\begin{equation*}
\left\langle\mathcal{H}_{v} \dot{\varphi}_{v}, \dot{\varphi}_{v}\right\rangle=-\ddot{\mathrm{d}}(v)<0 . \tag{4.6}
\end{equation*}
$$

Putting $y=a_{1} \chi+p_{1}, p_{1} \in P$, where $\chi$ is an eigenfunction of $\mathcal{H}_{v}$ corresponding to the negative eigenvalue $-\lambda_{0}^{2}$ and $P$ is the positive subspace of $\mathcal{H}_{v}$, we obtain

$$
\left\langle\mathcal{H}_{v} y, y\right\rangle=-a_{1}^{2} \lambda_{0}^{2}+\left\langle\mathcal{H}_{v} p_{1}, p_{1}\right\rangle .
$$

Write $\dot{\varphi}_{v}=a_{0} \chi+b_{0} \varphi_{v}^{\prime}+p_{0}, p_{0} \in P$. From (4.6) we have

$$
0>\left\langle\mathcal{H}_{v} \dot{\varphi}_{v}, \dot{\varphi}_{v}\right\rangle=-a_{0}^{2} \lambda_{0}^{2}+\left\langle\mathcal{H}_{v} p_{0}, p_{0}\right\rangle
$$

and

$$
\begin{aligned}
0 & =-\left\langle Q^{\prime}\left(\varphi_{v}\right), y\right\rangle=\left\langle H_{v} \dot{\varphi}_{v}, y\right\rangle \\
& =\left\langle-a_{0} \lambda_{0}^{2} \chi+H_{v} p_{0}, a_{1} \chi+p_{1}\right\rangle=-a_{0} a_{1} \lambda_{0}^{2}+\left\langle\mathcal{H}_{v} p_{0}, p_{1}\right\rangle .
\end{aligned}
$$

Hence

$$
\begin{aligned}
\left\langle\mathcal{H}_{v} y, y\right\rangle & =-a_{1}^{2} \lambda_{0}^{2}+\left\langle\mathcal{H}_{v} p_{1}, p_{1}\right\rangle \geqslant-a_{1}^{2} \lambda_{0}^{2}+\frac{\left\langle\mathcal{H}_{v} p_{0}, p_{1}\right\rangle^{2}}{\left\langle\mathcal{H}_{v} p_{0}, p_{0}\right\rangle} \\
& =-a_{1}^{2} \lambda_{0}^{2}+\frac{a_{0}^{2} a_{1}^{2} \lambda_{0}^{4}}{\lambda_{0}^{2} a_{0}^{2}-\ddot{\mathrm{d}}(v)}=a_{1}^{2} K_{1} .
\end{aligned}
$$

The rest of the proof is the same as in [33, pp 310-1].

Proposition 5. If assumption 1 holds, then for any positive $\varepsilon$ there exists $K>0$ such that for any solution of (1.1) satisfying $u \in \mathcal{U}_{\varepsilon}, Q(u)=Q\left(\varphi_{v}\right)$, one has

$$
\begin{equation*}
E(u)-E\left(\varphi_{v}\right)+c_{1} V(u)-c_{1} V\left(\varphi_{v}\right) \geqslant K\left\|u(.+\alpha(u))-\varphi_{v}\right\|_{1}^{2} . \tag{4.7}
\end{equation*}
$$

Proof. Denote $\psi=u(\cdot+\alpha(u))-\varphi_{v}=\mu\left(\varphi_{v}-\varphi_{v}^{\prime \prime}\right)+y$ where $\left\langle\varphi_{v}-\varphi_{v}^{\prime \prime}, y\right\rangle=0$. By the translation-invariant property of $Q$, we have

$$
\begin{aligned}
Q\left(\varphi_{v}\right) & =Q(u)=Q\left(\varphi_{v}\right)+\left\langle Q^{\prime}\left(\varphi_{v}\right), \psi\right\rangle+O\left(\|\psi\|_{1}^{2}\right) \\
& =Q\left(\varphi_{v}\right)+\left\langle\varphi_{v}-\varphi_{v}^{\prime \prime}, \mu\left(\varphi_{v}-\varphi_{v}^{\prime \prime}\right)+y\right\rangle+O\left(\|\psi\|_{1}^{2}\right) \\
& =Q\left(\varphi_{v}\right)+\mu\left\|\varphi_{v}-\varphi_{v}^{\prime \prime}\right\|_{L^{2}}^{2}+O\left(\|\psi\|_{1}^{2}\right)
\end{aligned}
$$

This implies $\mu=O\left(\|\psi\|_{1}^{2}\right)$. Since $Q(u)=Q\left(\varphi_{v}\right)$, the identity $M^{\prime \prime}\left(\varphi_{v}\right)=\mathcal{H}_{v}$ and the Taylor expansion give

$$
E(u)-E\left(\varphi_{v}\right)+c_{1} \int_{0}^{T} \psi \mathrm{~d} x=\frac{1}{2}\left\langle\mathcal{H}_{v} y, y\right\rangle+o\left(\|\psi\|_{1}^{2}\right) .
$$

From lemma 1 we have

$$
0=\left\langle u(\cdot+\alpha(u)), \varphi_{v}^{\prime}\right\rangle=\left\langle\mu\left(\varphi_{v}-\varphi_{v}^{\prime \prime}\right)+y+\varphi_{v}, \varphi_{v}^{\prime}\right\rangle=\left\langle y, \varphi_{v}^{\prime}\right\rangle .
$$

From the above equality and lemma 2 we obtain

$$
E(u)-E\left(\varphi_{v}\right)+c_{1} \int_{0}^{T} \psi \mathrm{~d} x \geqslant K\|y\|_{1}^{2}+o\left(\|\psi\|_{1}^{2}\right) .
$$

This estimate together with

$$
\|y\|_{1}=\left\|\psi-\mu\left(\varphi-\varphi^{\prime \prime}\right)\right\|_{1} \geqslant\|\psi\|_{1}-\left|\mu\left\|\mid \varphi-\varphi^{\prime \prime}\right\|_{1}\right.
$$

yield

$$
E(u)-E\left(\varphi_{v}\right)+c_{1} \int_{0}^{T} \psi \mathrm{~d} x \geqslant K\|\psi\|_{1}^{2}
$$

for $\|\psi\|_{1}$ sufficiently small. This completes the proof.
Theorem 2. If assumption 1 holds, then the travelling-wave solution $\varphi_{v}$ is stable .
Proof. Suppose that $\varphi_{v}$ is unstable. Then there exists a sequence of initial data $u_{n}(0) \in H^{1}$ and $\eta>0$ such that

$$
\left\|u_{n}(0)-\varphi_{v}\right\|_{1} \rightarrow 0
$$

but

$$
\begin{equation*}
\sup _{t \in\left[0, t_{0}\right]} \inf _{r \in \mathbb{R}}\left\|u_{n}(\cdot, t)-\varphi_{v}(\cdot-r)\right\|_{1} \geqslant \eta, \tag{4.8}
\end{equation*}
$$

where $u_{n} \in C\left(\left[0, t_{0}\right) ; H^{1}\right)$ is a solution of (1.1) with initial data $u_{n}(0)$. Let $t_{n} \in\left[0, t_{0}\right)$ be the first time so that

$$
\inf _{r \in \mathbb{R}}\left\|u_{n}\left(\cdot, t_{n}\right)-\varphi_{v}(\cdot-r)\right\|_{1}=\eta
$$

We have

$$
\begin{aligned}
& E\left(u_{n}\left(\cdot, t_{n}\right)\right)=E\left(u_{n}(0)\right) \rightarrow E\left(\varphi_{v}\right), \\
& Q\left(u_{n}\left(\cdot, t_{n}\right)\right)=Q\left(u_{n}(0)\right) \rightarrow Q\left(\varphi_{v}\right), \\
& V\left(u_{n}\left(\cdot, t_{n}\right)\right)=V\left(u_{n}(0)\right) \rightarrow V\left(\varphi_{v}\right) .
\end{aligned}
$$

Choose a sequence $\psi_{n} \in H^{1}$ such that $Q\left(\psi_{n}\right)=Q\left(\varphi_{v}\right)$ and $\left\|\psi_{n}-u_{n}\left(\cdot, t_{n}\right)\right\|_{1} \rightarrow 0$. By continuity of $E$ and $V, E\left(\psi_{n}\right) \rightarrow E\left(\varphi_{v}\right)$ and $V\left(\psi_{n}\right) \rightarrow V\left(\varphi_{v}\right)$. From (4.7) we have

$$
E\left(\psi_{n}\right)-E\left(\varphi_{v}\right)+c_{1}\left(V\left(\psi_{n}\right)-V\left(\varphi_{v}\right)\right) \geqslant K\left\|\psi_{n}\left(\cdot+\alpha\left(\psi_{n}\right)\right)-\varphi_{v}\right\|_{1}^{2} .
$$

Therefore $\left\|\psi_{n}-\varphi_{v}\left(\cdot-\alpha\left(\psi_{n}\right)\right)\right\|_{1} \rightarrow 0$, which implies

$$
\left\|u_{n}-\varphi_{v}\left(\cdot-\alpha\left(\psi_{n}\right)\right)\right\|_{1} \rightarrow 0
$$

This however contradicts (4.8). The proof of theorem 2 is complete.
In order to apply theorem 2 we have, according to assumption 1, to determine the sign of the derivative $\ddot{\mathrm{d}}(v)$. Using the same technique as in the proof of proposition 2, we obtain the following expression of $\ddot{\mathrm{d}}(v)$ through line integrals. When $a(u)$ and $b(u)$ are polynomials (or even rational functions), these are complete Abelian integrals. A lot of methods have been developed to investigate Abelian integrals, which could be applied here to study the sign of $\ddot{\mathrm{d}}(v)$.

Proposition 6. Assume that the period $T=T\left(\varphi_{0}\right)=\bar{T}(h)$ is not critical where $h$ is given by (3.11). Then

$$
\begin{equation*}
\ddot{\mathrm{d}}(v)=\frac{W(h)}{4 \bar{T}^{\prime}(h)}, \tag{4.9}
\end{equation*}
$$

with
$W(h)=\left(\frac{\mathrm{d}}{\mathrm{d} h} \oint_{H=h} \frac{\mathrm{~d} x}{y}\right)\left(\frac{\mathrm{d}}{\mathrm{d} h} \oint_{H=h} \frac{\left(x^{4}+2 x^{2} y^{2}-\frac{1}{3} y^{4}\right) \mathrm{d} x}{y}\right)-\left(\frac{\mathrm{d}}{\mathrm{d} h} \oint_{H=h} \frac{\left(x^{2}+y^{2}\right) \mathrm{d} x}{y}\right)^{2}$.
Proof. We apply again (3.13) and use the Gelfand-Leray form (in both directions) to calculate the needed derivatives. As

$$
\dot{\mathrm{d}}(v)=\frac{1}{2} \oint_{H=h} \frac{\left(x^{2}+y^{2}\right) \mathrm{d} x}{y}=\frac{1}{2} \oint_{H=h} y \mathrm{~d} x+\frac{1}{2} \frac{\mathrm{~d}}{\mathrm{~d} h} \oint_{H=h} x^{2}(v-b(x)) y \mathrm{~d} x
$$

one obtains

$$
\begin{aligned}
\ddot{\mathrm{d}}(v) & =\frac{\dot{h}}{2} \frac{\mathrm{~d}}{\mathrm{~d} h} \oint_{H=h} \frac{\left(x^{2}+y^{2}\right) \mathrm{d} x}{y}+\frac{1}{2} \frac{\mathrm{~d}}{\mathrm{~d} v}\left(\oint_{H=h} y \mathrm{~d} x+\frac{\mathrm{d}}{\mathrm{~d} h} \oint_{H=h} x^{2}(v-b(x)) y \mathrm{~d} x\right) \\
& =\frac{\dot{h}}{2} \frac{\mathrm{~d}}{\mathrm{~d} h} \oint_{H=h} \frac{\left(x^{2}+y^{2}\right) \mathrm{d} x}{y}+\oint_{H=h} \frac{\left(x^{2}-y^{2}\right) \mathrm{d} x}{4(v-b(x)) y}+\frac{\mathrm{d}}{\mathrm{~d} h} \oint_{H=h}\left(\frac{x^{2} y}{2}+\frac{x^{2}\left(x^{2}-y^{2}\right)}{4 y}\right) \mathrm{d} x \\
& =\frac{\dot{h}}{2} \frac{\mathrm{~d}}{\mathrm{~d} h} \oint_{H=h} \frac{\left(x^{2}+y^{2}\right) \mathrm{d} x}{y}+\frac{1}{4} \frac{\mathrm{~d}}{\mathrm{~d} h} \oint_{H=h} \frac{\left(x^{4}+2 x^{2} y^{2}-\frac{1}{3} y^{4}\right) \mathrm{d} x}{y} .
\end{aligned}
$$

Replacing the value of $\dot{h}$ from (3.16), we come to the needed formula.

## 5. Examples

Example 5.1 (the BBM equation). Consider the BBM equation

$$
\begin{equation*}
u_{t}+2 \omega u_{x}+3 u u_{x}-u_{x x t}=0, \quad \omega \in \mathbb{R} \tag{BBM}
\end{equation*}
$$

which is a particular case of (1.1) with $a(u)=2 \omega u+\frac{3}{2} u^{2}$ and $b=0$. To apply theorem 2 to BBM, we have to verify assumption 1 . Namely, to establish that the corresponding operator $\mathcal{H}_{v}$ has the needed spectral properties (i) and to prove the convexity of $\mathrm{d}(v)$.


Figure 1. BBM, case $v>0$.
Let us first mention that (3.5) reduces now to $D \equiv(v-2 \omega)^{2}+6 c_{1}>0$ and
$X_{0}=\frac{v-2 \omega+D^{1 / 2}}{3}, \quad \Delta=\left(\frac{v-2 \omega-D^{1 / 2}}{3}, \frac{v-2 \omega+2 D^{1 / 2}}{3}\right) \quad$ if $\quad v>0$,
$X_{0}=\frac{v-2 \omega-D^{1 / 2}}{3}, \quad \Delta=\left(\frac{v-2 \omega-2 D^{1 / 2}}{3}, \frac{v-2 \omega+D^{1 / 2}}{3}\right) \quad$ if $\quad v<0$.
By the definition of $c_{1}, c_{2}$ and $U(s)$, one obtains in the considered case

$$
\begin{aligned}
U(s) & \equiv \frac{1}{v}\left(\varphi_{0}-s\right)\left[s^{2}+\left(\varphi_{0}+2 \omega-v\right) s-\left(2 \varphi_{0}^{2}+(2 \omega-v) \varphi_{0}+2 v \varphi_{2}\right)\right] \\
& =\frac{\left(s-\varphi_{0}\right)\left(\varphi_{1}-s\right)\left(s+\varphi_{1}+\varphi_{0}+2 \omega-v\right)}{v}
\end{aligned}
$$

We note that the last equality is a consequence of proposition 1 (ii) which implies that $U\left(\varphi_{1}\right)=U\left(\varphi_{0}\right)=0$. To obtain an explicit formula for the travelling wave $\varphi_{v}$, we substitute $\sigma=\varphi_{0}+\left(\varphi_{1}-\varphi_{0}\right) z^{2}, z>0$, in order to express the integral in (3.6) as an elliptic integral of the first kind in a Legendre form. If $v<0$, one obtains

$$
\int_{0}^{Z(s)} \frac{\mathrm{d} z}{\sqrt{\left(1-z^{2}\right)\left(1-k^{2} z^{2}\right)}}=\alpha(s-\xi)
$$

where
$Z(s)=\sqrt{\frac{\varphi_{v}(s)-\varphi_{0}}{\varphi_{1}-\varphi_{0}}}, \quad k^{2}=-\frac{\varphi_{1}-\varphi_{0}}{\varphi_{1}+2 \varphi_{0}+2 \omega-v}, \quad \alpha=\sqrt{\frac{\varphi_{1}+2 \varphi_{0}+2 \omega-v}{4 v}}$.
Thus, we get the expression

$$
\begin{equation*}
\varphi_{v}(s)=\varphi_{0}+\left(\varphi_{1}-\varphi_{0}\right) s n^{2}(\alpha(s-\xi) ; k) . \tag{5.1}
\end{equation*}
$$

Similarly, in the case $v>0$ (see figure 1) we obtain (with the same $Z$ )

$$
\int_{0}^{Z(s)} \frac{\mathrm{d} z}{\sqrt{\left(1-z^{2}\right)\left(k^{\prime 2}+k^{2} z^{2}\right)}}=\alpha(s-\xi)
$$

where
$k^{2}=\frac{\varphi_{1}-\varphi_{0}}{\varphi_{0}+2 \varphi_{1}+2 \omega-v}, \quad k^{2}+k^{\prime 2}=1, \quad \alpha=\sqrt{\frac{\varphi_{0}+2 \varphi_{1}+2 \omega-v}{4 v}}$,
and the expression for $\varphi_{v}$

$$
\begin{equation*}
\varphi_{v}(s)=\varphi_{0}+\left(\varphi_{1}-\varphi_{0}\right) c n^{2}(\alpha(s-\xi) ; k) . \tag{5.2}
\end{equation*}
$$

To calculate the period of $\varphi_{v}$, we use (3.7) and the same procedure as above. In this way we get in both cases

$$
T=2 \int_{\varphi_{0}}^{\varphi_{1}} \frac{\mathrm{~d} \sigma}{\sqrt{U(\sigma)}}=\frac{2}{\alpha} \int_{0}^{1} \frac{\mathrm{~d} z}{\sqrt{\left(1-z^{2}\right)\left(1-k^{2} z^{2}\right)}}=\frac{2 K(k)}{\alpha} .
$$

We return to the operator $\mathcal{H}_{v}$ defined by (4.4) which now has the form

$$
\begin{equation*}
\mathcal{H}_{v}=-v \partial_{x}^{2}+v-2 \omega-3 \varphi_{v} \tag{5.3}
\end{equation*}
$$

where $\varphi_{v}$ is determined by (5.1) or (5.2). Take $v>0$ and consider the spectral problem

$$
\begin{equation*}
\mathcal{H}_{v} \psi=\lambda \psi, \quad \psi(0)=\psi(T), \quad \psi^{\prime}(0)=\psi^{\prime}(T) \tag{5.4}
\end{equation*}
$$

We will denote the operator just defined again by $\mathcal{H}_{v}$. It is a self-adjoint operator acting in $H^{2}([0, T])$. From the Floquet theory applied to (5.4) it follows [44] that its spectrum is purely discrete,

$$
\begin{equation*}
\lambda_{0}<\lambda_{1} \leqslant \lambda_{2}<\lambda_{3} \leqslant \lambda_{4}<\cdots \tag{5.5}
\end{equation*}
$$

where $\lambda_{0}$ is always a simple eigenvalue. If $\psi_{n}(x)$ is the eigenfunction corresponding to $\lambda_{n}$, then

$$
\begin{align*}
& \psi_{0} \text { has no zeros in }[0, T] ; \\
& \psi_{2 n+1}, \quad \psi_{2 n+2} \text { have each just } 2 n+2 \text { zeros in }[0, T) \tag{5.6}
\end{align*}
$$

Proposition 7. The linear operator $\mathcal{H}_{v}$ defined by (5.3) and (5.4) has the following spectral properties for $v>0$ :
(i) The first three eigenvalues of $\mathcal{H}_{v}$ are simple.
(ii) The second eigenvalue of $\mathcal{H}_{v}$ is $\lambda_{1}=0$.

Proof. By (3.1), $\mathcal{H}_{v} \varphi_{v}^{\prime}=0$, hence $\psi=\varphi_{v}^{\prime}$ is an eigenfunction corresponding to zero eigenvalue. By proposition 1(iii), $\varphi^{\prime}$ has just two zeros in [0,T) and therefore by (5.6) either $0=\lambda_{1}<\lambda_{2}$ or $\lambda_{1}<\lambda_{2}=0$ or $\lambda_{1}=\lambda_{2}=0$. We are going to verify that only the first possibility $0=\lambda_{1}<\lambda_{2}$ can occur. From the definition of $k$ and $\alpha$, one obtains that

$$
\varphi_{0}+2 \varphi_{1}+2 \omega-v=4 v \alpha^{2}, \quad \varphi_{1}-\varphi_{0}=4 v k^{2} \alpha^{2}
$$

Then using (5.2) we get

$$
\begin{aligned}
\mathcal{H}_{v} & =-v \partial_{x}^{2}+v-2 \omega-3 \varphi_{0}-3\left(\varphi_{1}-\varphi_{0}\right) c n^{2}(\alpha x ; k) \\
& =-v \partial_{x}^{2}+v-2 \omega-3 \varphi_{1}+3\left(\varphi_{1}-\varphi_{0}\right) s n^{2}(\alpha x ; k) \\
& =-v \partial_{x}^{2}-v \alpha^{2}\left[4 k^{2}+4-12 k^{2} \operatorname{sn}^{2}(\alpha x ; k)\right] \\
& =v \alpha^{2}\left[-\partial_{y}^{2}-4 k^{2}-4+12 k^{2} \operatorname{sn}^{2}(y ; k)\right] \equiv v \alpha^{2} \Lambda
\end{aligned}
$$

where $y=\alpha x$. The operator $\Lambda$ is related to Hill's equation with Lamé potential

$$
\Lambda w=-\frac{\mathrm{d}^{2}}{\mathrm{~d} y^{2}} w+\left[12 k^{2} \operatorname{sn}^{2}(y ; k)-4 k^{2}-4\right] w=0
$$

and its spectral properties in the interval $[0,2 K(k)]$ are well known $[2,5,36]$. The first three (simple) eigenvalues and corresponding periodic eigenfunctions of $\Lambda$ are

$$
\begin{aligned}
& \mu_{0}=k^{2}-2-2 \sqrt{1-k^{2}+4 k^{4}}<0, \\
& \psi_{0}(y)=\mathrm{d} n(y ; k)\left[1-\left(1+2 k^{2}-\sqrt{1-k^{2}+4 k^{4}}\right) s n^{2}(y ; k)\right]>0, \\
& \mu_{1}=0 \\
& \psi_{1}(y)=\mathrm{d} n(y ; k) \operatorname{sn}(y ; k) c n(y ; k)=\frac{1}{2}(d / \mathrm{d} y) c n^{2}(y ; k), \\
& \mu_{2}=k^{2}-2+2 \sqrt{1-k^{2}+4 k^{4}}>0, \\
& \psi_{2}(y)=\mathrm{d} n(y ; k)\left[1-\left(1+2 k^{2}+\sqrt{1-k^{2}+4 k^{4}}\right) s n^{2}(y ; k)\right] .
\end{aligned}
$$

As the eigenvalues of $\mathcal{H}_{v}$ and $\Lambda$ are related by $\lambda_{n}=v \alpha^{2} \mu_{n}$, we conclude that for $v>0$ the first three eigenvalues of (5.3) and (5.4) are simple and moreover $\lambda_{0}<0, \lambda_{1}=0, \lambda_{2}>0$. The corresponding eigenfunctions are $\psi_{0}(\alpha x), \psi_{1}(\alpha x)=$ const $\cdot \varphi_{v}^{\prime}(x)$ and $\psi_{2}(\alpha x)$.


Figure 2. Three topologically different phase portraits of the three cases related to the symmetric Hamiltonian.

What about the sign of $\ddot{\mathrm{d}}(v)$, it is easily seen it is positive when $\omega=0$ (see the end of this section). For $\omega \neq 0$, the proof is much more long and complicated. It can be done by following the procedure we propose in the following subsection. For this reason, we are not going to investigate the case $\omega \neq 0$ in the present paper.

Example 5.2 (the modified BBM equation). Our second example is concerned with the periodic travelling-wave solutions of the equation

$$
\begin{equation*}
u_{t}+2 w u_{x}+\beta\left(u^{3}\right)_{x}-u_{x x t}=0 \tag{mBBM}
\end{equation*}
$$

where $\omega, \beta \in \mathbb{R}$ and $\beta \neq 0$. For this case, $a(u)=2 w u+\beta u^{3}$ and $b(u)=0$ in (1.1). For definiteness, we take $\varphi_{v}=\varphi(x-v t)$ where $v>0$. The Hamiltonian flow in (3.4) is generated by

$$
H(X, Y)=\frac{v}{2} Y^{2}+\frac{\beta}{4} X^{4}+\left(\omega-\frac{v}{2}\right) X^{2}-c_{1} X
$$

Our plan is to study here the 'symmetric' case $c_{1}=0$. The general case could then be considered as a perturbation of the symmetric one (at least for $c_{1}$ small). It is well known that there are three cases related to the symmetric Hamiltonian

$$
Y^{2}+\frac{\beta}{2 v} X^{4}+\left(\frac{2 \omega}{v}-1\right) X^{2}, \quad v>0
$$

see, e.g. [35]:
(i) global centre: $\beta>0,2 \omega>v$;
(ii) truncated pendulum: $\beta<0,2 \omega>v$;
(iii) Duffing oscillator: $\beta>0,2 \omega<v$,
with three topologically different phase portraits (see figure 2). There is one continuous family of periodic orbits in cases (i) and (ii) and three families (left, right, outer) in (iii):

This also holds true for all $c_{1}$ in case (i) and for $c_{1}^{2}<\frac{4}{27} \beta^{-1}(v-2 \omega)^{3}$ in (ii) and (iii). In the symmetric case $c_{1}=0$ we deal with, the periodic solutions exist if and only if $\varphi_{0} \in \Delta^{-}$ (see proposition 1), where $\Delta^{-}=(-\infty, 0)$ in case (i); $\Delta^{-}=\left(-\left(\frac{v-2 \omega}{\beta}\right)^{1 / 2}, 0\right)$ in case (ii); $\Delta^{-}=\left(-\left(\frac{2 v-4 \omega}{\beta}\right)^{1 / 2},-\left(\frac{v-2 \omega}{\beta}\right)^{1 / 2}\right)$ in case (iii), left family; $\Delta^{-}=\left(0,\left(\frac{v-2 \omega}{\beta}\right)^{1 / 2}\right)$ in case (iii), right family; $\Delta^{-}=\left(-\infty,-\left(\frac{2 v-4 \omega}{\beta}\right)^{1 / 2}\right)$ in case (iii), outer family. The respective functions
$U(s)$ take the form:
$U(s)=\frac{\beta}{2 v}\left(\varphi_{0}^{2}-s^{2}\right)\left(\frac{4 \omega-2 v}{\beta}+\varphi_{0}^{2}+s^{2}\right) \quad$ in case (i) and case (iii), outer family,
$U(s)=-\frac{\beta}{2 v}\left(\varphi_{0}^{2}-s^{2}\right)\left(\frac{2 v-4 \omega}{\beta}-\varphi_{0}^{2}-s^{2}\right) \quad$ in case (ii),
$U(s)=\frac{\beta}{2 v}\left(s^{2}-\varphi_{1}^{2}\right)\left(\varphi_{0}^{2}-s^{2}\right) \quad$ in case (iii), left family,
$U(s)=\frac{\beta}{2 v}\left(s^{2}-\varphi_{0}^{2}\right)\left(\varphi_{1}^{2}-s^{2}\right) \quad$ in case (iii), right family.
In a similar way as we have done in the (BBM) case, we can use formula (3.6) to calculate explicitly $\varphi_{v}$. By an appropriate change of the variables in (3.6), one can express the solution through standard elliptic integrals (we omit the details). Up to a translation of the argument, we have
$\varphi_{v}(s)=\varphi_{0} c n(\alpha s ; k), \quad \alpha=\sqrt{\frac{2 \omega-v+\beta \varphi_{0}^{2}}{v}}, \quad k=\sqrt{\frac{\beta \varphi_{0}^{2}}{4 \omega-2 v+2 \beta \varphi_{0}^{2}}}$, (i) and (iii) outer,
$\varphi_{v}(s)=\varphi_{0} \operatorname{sn}(\alpha s ; k), \quad \alpha=\sqrt{\frac{4 \omega-2 v+\beta \varphi_{0}^{2}}{2 v}}, \quad k=\sqrt{\frac{\beta \varphi_{0}^{2}}{2 v-4 \omega-\beta \varphi_{0}^{2}}}$,
$\varphi_{v}(s)=\varphi_{0} \mathrm{~d} n(\alpha s ; k), \quad \alpha=\sqrt{\frac{\beta \varphi_{0}^{2}}{2 v}}, \quad k=\sqrt{\frac{4 \omega-2 v+2 \beta \varphi_{0}^{2}}{\beta \varphi_{0}^{2}}}, \quad$ (iii) left,
$\varphi_{v}(s)=\varphi_{1} \mathrm{~d} n(\alpha s ; k), \quad \alpha=\sqrt{\frac{\beta \varphi_{1}^{2}}{2 v}}, \quad k=\sqrt{\frac{4 \omega-2 v+2 \beta \varphi_{1}^{2}}{\beta \varphi_{1}^{2}}}, \quad$ (iii) right.
These formulae of $\alpha$ and $k$ yield the following expressions and range $I$ for the period
$T=4 \sqrt{\frac{v}{2 \omega-v}} \sqrt{1-2 k^{2}} K(k), \quad k \in\left(0, \frac{1}{\sqrt{2}}\right), \quad I=\left(0,2 \pi \sqrt{\frac{v}{2 \omega-v}}\right)$,
$T=4 \sqrt{\frac{v}{2 \omega-v}} \sqrt{1+k^{2}} K(k), \quad k \in(0,1), \quad I=\left(2 \pi \sqrt{\frac{v}{2 \omega-v}}, \infty\right), \quad$ (ii),
$T=2 \sqrt{\frac{v}{v-2 \omega}} \sqrt{2-k^{2}} K(k), \quad k \in(0,1), \quad I=\left(2 \pi \sqrt{\frac{v}{2 v-4 \omega}}, \infty\right), \quad$ (iii) left and right,
$T=4 \sqrt{\frac{v}{v-2 \omega}} \sqrt{2 k^{2}-1} K(k), \quad k \in\left(\frac{1}{\sqrt{2}}, 1\right), \quad I=(0, \infty), \quad$ (iii) outer.
Indeed, the formulae just derived imply that $T=T(k)$ is strictly decreasing in case (i) and strictly increasing in the other cases. In fact, for (i), (ii) and (iii) outer this follows already from (3.11) and [28]. As far as cases (iii) left and right are concerned, now the result follows from

$$
\frac{\mathrm{d}}{\mathrm{~d} k}\left(\sqrt{2-k^{2}} K(k)\right)=\frac{\left(2-k^{2}\right) K^{\prime}-k K}{\sqrt{2-k^{2}}}=\frac{K^{\prime}+E^{\prime}}{\sqrt{2-k^{2}}}>0 .
$$

On the other hand, in all cases one has $\mathrm{d} k / \mathrm{d} \varphi_{0} \neq 0$. Therefore, given $T \in I$, the condition holds in order to determine $\varphi_{0}$ by the implicit function theorem so that the respective $\varphi$ would have a period $T$.

Finally, by using the above formulae one easily obtains that

$$
\mathcal{H}_{v}=-v \partial_{x}^{2}+v-2 \omega-3 \beta \varphi_{v}^{2}=v \alpha^{2}\left[-\partial_{y}^{2}+6 k^{2} s n^{2}(y ; k)+m\right]
$$

where $y=\alpha x$ and $m=-1-4 k^{2}$ in cases (i) and (iii) outer, $m=-4-k^{2}$ in cases (iii) left and right, $m=-1-k^{2}$ in case (ii), respectively.

Lemma 3. The first five eigenvalues of the operator $\Lambda$ defined by the differential expression $\Lambda=-\partial_{y}^{2}+6 k^{2} s n^{2}(y ; k)$, with periodic boundary conditions on $[0,4 K(k)]$, are simple. These eigenvalues and their respective eigenfunctions are
$\mu_{0}=2+2 k^{2}-2 \sqrt{1-k^{2}+k^{4}}, \quad \psi_{0}(y)=1-\left(1+k^{2}-\sqrt{1-k^{2}+k^{4}}\right) \operatorname{sn}^{2}(y ; k)$,
$\mu_{1}=1+k^{2}, \quad \psi_{1}(y)=c n(y ; k) \mathrm{d} n(y ; k)=s n^{\prime}(y ; k)$,
$\mu_{2}=1+4 k^{2}, \quad \psi_{2}(y)=\operatorname{sn}(y ; k) \mathrm{d} n(y ; k)=-c n^{\prime}(y ; k)$,
$\mu_{3}=4+k^{2}, \quad \psi_{3}(y)=\operatorname{sn}(y ; k) c n(y ; k)=-k^{-2} \mathrm{~d} n^{\prime}(y ; k)$,
$\mu_{4}=2+2 k^{2}+2 \sqrt{1-k^{2}+k^{4}}, \quad \psi_{4}(y)=1-\left(1+k^{2}+\sqrt{1-k^{2}+k^{4}}\right) \operatorname{sn}^{2}(y ; k)$.

Proof. The equalities $\Lambda \psi_{n}(y)=\mu_{n} \psi_{n}(y), 0 \leqslant n \leqslant 4$ are established by calculation. By (5.5), (5.6) and the properties of elliptic functions, $\mu_{n}$ are simple and the rest of the spectrum lies in the interval $\left(\mu_{4}, \infty\right)$.

Corollary 1. The first three eigenvalues of the operator $\Lambda$, equipped with periodic boundary conditions on $[0,2 K(k)]$, are simple and equal to $\mu_{0}, \mu_{3}, \mu_{4}$.

By lemma 3 and its corollary, $\varphi_{v}^{\prime}$ is the second eigenfunction of the operator $\mathcal{H}_{v}$ in the cases: truncated pendulum and Duffing oscillator (left and right). For these cases, it makes sense to investigate the sign of $\ddot{d}(v)$ which we do in the following proposition.

Proposition 8. Assume that $T \in I=\left(T_{0}, \infty\right)$ and $\varphi_{v}=\varphi(x-v t), v>0$, is the periodic travelling-wave solution having a minimal period $T$. Then
(i) in the truncated pendulum case, one has $\ddot{\mathrm{d}}(v)<0$;
(ii) in the left (right) Duffing oscillator case, if $3 v^{2}-8 \omega^{2} \geqslant 0$, then $\ddot{\mathrm{d}}(v)>0$. If $3 v^{2}-8 \omega^{2}<0$ and $2 v^{2}-2 \omega v-\omega^{2}>0$, then there is $T_{\max } \in\left(T_{0}, \infty\right)$ depending only on the ratio $\omega / v$, such that $\ddot{\mathrm{d}}(v)>0$ for $T>T_{\max }$ and $\ddot{\mathrm{d}}(v)<0$ for $T \in\left(T_{0}, T_{\max }\right)$. If $2 v^{2}-2 \omega v-\omega^{2} \leqslant 0$, then $\ddot{\mathrm{d}}(v)<0$.

Proof. We are going to use formula (4.9). In the example we deal with, one has

$$
y^{2}=U(x, h)=\frac{2 h}{v}-\frac{2 \omega-v}{v} x^{2}-\frac{\beta}{2 v} x^{4}, \quad h=\frac{2 \omega-v}{2} \varphi_{0}^{2}+\frac{\beta}{4} \varphi_{0}^{4} .
$$

Given a non-negative even integer $n$, denote $I_{n}(h)=\oint_{H=h} x^{n} y \mathrm{~d} x$. It is well known that the linear space of integrals $\left\{I_{n}(h), n\right.$ even $\}$ forms a polynomial $R[h]$ module with two generators, $I_{0}(h)$ and $I_{2}(h)$. Moreover, $I_{0}$ and $I_{2}$ satisfy a Picard-Fuchs system of dimension two. This implies that the ratio $R(h)=I_{2}^{\prime}(h) / I_{0}^{\prime}(h)$ satisfies a Riccati equation. We shall use these facts to express $\ddot{\mathrm{d}}(v)$ as a quadratic form with respect to $I_{0}^{\prime}, I_{2}^{\prime}$ with polynomial coefficients in $h$ and use the properties of the Riccati equation to determine the sign of $\ddot{\mathrm{d}}(v)$. The procedure might seem too long, but it is universal (at least in the case when $a(u)$ and $b(u)$ are polynomials) and therefore applicable to many other cases.

Let us first express $\ddot{\mathrm{d}}(v)$ through the integrals $I_{n}$. Below, we will denote for short the derivatives with respect to $h$ by $I_{n}^{\prime}, I_{n}^{\prime \prime}$, etc. Using the first equality in (3.13) with $b=0$, we obtain

$$
I_{n}^{\prime}(h)=\oint_{H=h} \frac{x^{n} \mathrm{~d} x}{v y}
$$

On the other hand,

$$
\oint_{H=h} y^{3} \mathrm{~d} x=\oint_{H=h} U(x, h) y \mathrm{~d} x=\frac{2 h}{v} I_{0}-\frac{2 \omega-v}{v} I_{2}-\frac{\beta}{2 v} I_{4} .
$$

By using these expressions, we obtain that
$\ddot{\mathrm{d}}(v)=\frac{v}{4 I_{0}^{\prime \prime}}\left[I_{0}^{\prime \prime}\left(I_{4}^{\prime}+\frac{\beta}{6 v^{2}} I_{4}+\frac{2 \omega+5 v}{3 v^{2}} I_{2}-\frac{2 h}{3 v^{2}} I_{0}\right)^{\prime}-\left(\frac{1}{v} I_{0}^{\prime}+I_{2}^{\prime \prime}\right)^{2}\right]$.
For reader's convenience, below we proceed to derive the relations between integrals $I_{n}$ and the Picard-Fuchs system satisfied by $I_{0}$ and $I_{2}$.

Lemma 4. (i) The following relations hold:

$$
\begin{equation*}
(n+6) \beta I_{n+3}+(2 n+6)(2 \omega-v) I_{n+1}=4 n h I_{n-1}, \quad n=1,3,5, \ldots \tag{5.8}
\end{equation*}
$$

(ii) The integrals $I_{0}$ and $I_{2}$ satisfy the system
$4 h I_{0}^{\prime}-(2 \omega-v) I_{2}^{\prime}=3 I_{0}, \quad-\frac{4(2 \omega-v)}{3 \beta} h I_{0}^{\prime}+\left(4 h+\frac{4(2 \omega-v)^{2}}{3 \beta}\right) I_{2}^{\prime}=5 I_{2}$.
Proof. (i). Integrating by parts, we obtain the identity

$$
\begin{equation*}
\oint_{H=h}\left[x^{n} U^{\prime}(x)+\frac{2}{3} n x^{n-1} U(x)\right] y \mathrm{~d} x=0 . \tag{5.10}
\end{equation*}
$$

Indeed,

$$
\begin{gathered}
\oint_{H=h} x^{n} U^{\prime}(x) y \mathrm{~d} x=\oint_{H=h} x^{n} y \mathrm{~d} y^{2}=\frac{2}{3} \oint_{H=h} x^{n} \mathrm{~d} y^{3}=-\frac{2}{3} n \oint_{H=h} x^{n-1} y^{3} \mathrm{~d} x \\
=-\frac{2}{3} n \oint_{H=h} x^{n-1} U(x) y \mathrm{~d} x
\end{gathered}
$$

Identity (5.10) is equivalent to (5.8).
(ii) Similarly, one has
$\oint_{H=h} \frac{x^{n} U^{\prime}(x) \mathrm{d} x}{y}=\oint_{H=h} \frac{x^{n} \mathrm{~d} y^{2}}{y}=\oint_{H=h} 2 x^{n} \mathrm{~d} y=-2 n \oint_{H=h} x^{n-1} y \mathrm{~d} x$,
$\oint_{H=h} \frac{x^{n} U(x) \mathrm{d} x}{y}=\oint_{H=h} x^{n} y \mathrm{~d} x$.
Rewriting these identities by means of $I_{n}$, we come to the formulae
$\beta I_{n+3}^{\prime}+(2 \omega-v) I_{n+1}^{\prime}=n I_{n-1}, \quad-\beta I_{n+4}^{\prime}-2(2 \omega-v) I_{n+2}^{\prime}+4 h I_{n}^{\prime}=2 I_{n}$.
The last two relations imply

$$
\begin{equation*}
4 h I_{n}^{\prime}-(2 \omega-v) I_{n+2}^{\prime}=(n+3) I_{n} \tag{5.13}
\end{equation*}
$$

Using (5.13) with $n=0$, 2 , we obtain the system

$$
4 h I_{0}^{\prime}-(2 \omega-v) I_{2}^{\prime}=3 I_{0}, \quad 4 h I_{2}^{\prime}-(2 \omega-v) I_{4}^{\prime}=5 I_{2}
$$

We remove $I_{4}^{\prime}$ from the last equation by using the first equation in (5.12). As a result one obtains (5.9).

After proving lemma 4 , we return to our main goal. Using the identities

$$
I_{4}=\frac{4 h}{7 \beta} I_{0}-\frac{8(2 \omega-v)}{7 \beta} I_{2}, \quad I_{4}^{\prime}=\frac{1}{\beta} I_{0}-\frac{2 \omega-v}{\beta} I_{2}^{\prime}
$$

and the first equation in (5.9), we remove $I_{4}, I_{4}^{\prime}$ and $I_{0}$ from (5.7). The result is
$\ddot{\mathrm{d}}(v)=\frac{v}{4 I_{0}^{\prime \prime}}\left[I_{0}^{\prime \prime}\left(\frac{3 v^{2}-4 \beta h}{3 \beta v^{2}} I_{0}^{\prime}+\frac{2 \omega+5 v}{3 v^{2}} I_{2}^{\prime}-\frac{2 \omega-v}{\beta} I_{2}^{\prime \prime}\right)-\left(\frac{1}{v} I_{0}^{\prime}+I_{2}^{\prime \prime}\right)^{2}\right]$.
Now, we differentiate (5.9) and determine the second derivatives from the obtained system:

$$
\begin{align*}
& \mathcal{D}(h) I_{0}^{\prime \prime}=-4 h I_{0}^{\prime}+(2 \omega-v) I_{2}^{\prime} \\
& \mathcal{D}(h) I_{2}^{\prime \prime}=\frac{4(2 \omega-v)}{\beta} h I_{0}^{\prime}+4 h I_{2}^{\prime}  \tag{5.15}\\
& \mathcal{D}(h)=16 h^{2}+\frac{4(2 \omega-v)^{2}}{\beta} h
\end{align*}
$$

Replacing in (5.14) and performing some direct calculations to simplify the result, we derive the formula
$\ddot{\mathrm{d}}(v)=\frac{1}{9 v I_{0}}\left[\left(8 h^{2}+\frac{12 \omega^{2}}{\beta} h\right) I_{0}^{\prime 2}+(4 \omega+10 v) h I_{0}^{\prime} I_{2}^{\prime}+\left(2 v^{2}-2 \omega v-\omega^{2}\right) I_{2}^{\prime 2}\right]$.
By (5.15), the ratio $R(h)=I_{2}^{\prime}(h) / I_{0}^{\prime}(h)$ satisfies the Riccati equation

$$
\begin{equation*}
\mathcal{D}(h) R^{\prime}=\frac{4(2 \omega-v)}{\beta} h+8 h R-(2 \omega-v) R^{2} \tag{5.16}
\end{equation*}
$$

In order to remove the parameters from (5.16), we take new variables $\bar{h}, \bar{R}$ through

$$
h=-\frac{(2 \omega-v)^{2}}{8 \beta}(\bar{h}+1), \quad R(h)=-\frac{2 \omega-v}{2 \beta}(\bar{R}(\bar{h})+1) .
$$

Thus, we obtain the final formulae

$$
\ddot{\mathrm{d}}(v)=\frac{(2 \omega-v)^{2} I_{0}^{\prime 2}(h)}{72 v \beta^{2} I_{0}(h)} w(\bar{h}, \bar{R}),
$$

where (below we will omit thoroughly the bars)

$$
\begin{gather*}
w(h, R)=\left(4 v^{2}-4 \omega v-2 \omega^{2}\right) R^{2}+\left(4 \omega^{2}+8 \omega v-5 v^{2}\right) h R \\
+(2 \omega-v)^{2} h^{2}+3 v^{2}(R-h)-6 \omega^{2} \tag{5.17}
\end{gather*}
$$

and $R=R(h)$ satisfies the equation (and related system with respect to a dummy variable)

$$
4\left(1-h^{2}\right) R^{\prime}=1-2 h R+R^{2}, \quad \begin{array}{ll}
\dot{h}=4\left(1-h^{2}\right)  \tag{5.18}\\
& \dot{R}=1-2 h R+R^{2}
\end{array}
$$



Figure 3. Phase portraint of (5.18).

Since $v>0$ and $I_{0}>0$ (this is the area integral), the sign of $\ddot{\mathrm{d}}(v)$ is determined by $w$. The conic curve $w(h, R)=0$ divides the $\mathbb{R}^{2}$ plane into two parts $W_{+}$and $W_{-}$according to the sign of $w(h, R)$. We have to identify the curve $\Gamma$ in the phase portrait of (5.18) (see figure 3) which corresponds to $I_{2}^{\prime} / I_{0}^{\prime}$ and determine its location with respect to $W_{+}$and $W_{-}$. In the case when $\ddot{\mathrm{d}}(v)$ changes sign, the main problem will be to prove that the conic curve intersects $\Gamma$ at most once.

It is easy to see that system (5.18) has two critical points $(1,1)$ and $(-1,-1)$ which are saddle nodes. They are connected by two separatrix trajectories, upper $\Gamma_{u}$ and lower $\Gamma_{l}$, see picture 3. Obviously, the phase portrait of (5.18) is symmetric with respect to the origin.
Lemma 5. In (5.18), the trajectory corresponding to the truncated pendulum case is $\Gamma_{l}$. The trajectory corresponding to the left and right Duffing oscillator cases is $\Gamma_{u}$.

Proof. If $\varphi_{0} \in \Delta_{-}$, then $\varphi_{0}^{2} \in\left(0, \frac{v-2 \omega}{\beta}\right)$ in the truncated pendulum and right Duffing oscillator cases and $\varphi_{0}^{2} \in\left(\frac{v-2 \omega}{\beta}, \frac{2 v-4 \omega}{\beta}\right)$ in the left case. Correspondingly, one obtains

$$
\begin{aligned}
& h \in\left(0,-\frac{(2 \omega-v)^{2}}{4 \beta}\right)=\left(h_{c}, h_{s}\right) \text { in the truncated pendulum case, } \\
& h \in\left(-\frac{(2 \omega-v)^{2}}{4 \beta}, 0\right)=\left(h_{c}, h_{s}\right) \text { in the left and right Duffing oscillator cases, }
\end{aligned}
$$

where $h_{c}$ is the Hamiltonian level corresponding to a centre and $h_{s}$ to a saddle. Moreover, $R(h)=I_{2}^{\prime}(h) / I_{0}^{\prime}(h)$ is analytic in a neighbourhood of $h=h_{c}$ and, by the mean value theorem,

$$
\lim _{h \rightarrow h_{c}} R(h)=\lim _{h \rightarrow h_{c}} \frac{I_{2}(h)}{I_{0}(h)}=X_{0}^{2} .
$$

We recall that $X_{0}$ is the abscissa of the centre and $X_{0}=0$ in the truncated pendulum case, $X_{0}^{2}=\frac{v-2 \omega}{\beta}$ in the left and right Duffing oscillator case. On its turn, by Picard-Lefschetz theory [6, chapter 3], near $h=h_{s}$ the integral $I_{n}(h)$ has the expansion $I_{n}(h)=I_{n}\left(h_{s}\right)+\alpha_{n}\left(h-h_{s}\right) \log \left|h-h_{s}\right|+\beta_{n}\left(h-h_{s}\right)+\gamma_{n}\left(h-h_{s}\right)^{2} \log \left|h-h_{s}\right|+\cdots$,
where $\alpha_{0} \neq 0$. Therefore, $R(h)=I_{2}^{\prime}(h) / I_{0}^{\prime}(h)$ is bounded near $h=h_{s}$. In terms of the new coordinates ( $\bar{h}, \bar{R}$ ), all this means that the phase trajectory of (5.18) corresponding to the
truncated pendulum case is $(h, R(h))$ where $-1<h<1, R(h) \rightarrow \pm 1$ as $h \rightarrow \pm 1$, and $R(h)$ is analytic near $h=-1$. The unique phase curve with these properties is $\Gamma_{l}$. Similarly, the phase trajectory of (5.18) corresponding to the left and right Duffing oscillator cases has the same ends but is analytic near $h=1$, hence it is $\Gamma_{u}$.

Below, we list some properties of the separatrices $\Gamma_{l}, \Gamma_{u}$ and the conic curve $w=0$. Apart from the separatrices, the conic curve depends on one parameter $\omega / v$ and undergoes several bifurcations when $\omega / v$ varies.

Lemma 6. The separatrices $\Gamma_{u}$ and $\Gamma_{l}$ have the following properties:
(i) $\Gamma_{u}$ is increasing and concave, $\Gamma_{l}$ is increasing and convex.
(ii) The tangent to $\Gamma_{u}$ at $(-1,-1)$ is $h=-1$, the tangent to $\Gamma_{l}$ at $(1,1)$ is $h=1$.
(iii) The tangent to $\Gamma_{u}$ at $(1,1)$ and the tangent to $\Gamma_{l}$ at $(-1,-1)$ have a slope $\frac{1}{4}$.

Proof. The proof easily follows from (5.18). Claim (iii) is a consequence of the analyticity at the corresponding point. To establish (ii), we use the expansions containing logarithmic terms above. They imply that the asymptotic expansion near $(-1,-1)$ of the non-analytic trajectories has the form $-1+\mu / \log (h+1)+\cdots$. We determine $\mu$ from (5.18) and obtain that on $\Gamma_{u}$ one has

$$
\begin{equation*}
R(h)=-1-\frac{8}{\log (h+1)}+O\left(\log ^{-2}(h+1)\right) \tag{5.19}
\end{equation*}
$$

A similar formula holds for $\Gamma_{l}$, hence (ii) is proved. Finally, in the strip $|h|<1$ one has $1-2 h R+R^{2}>0$, therefore $R^{\prime}>0$ and all trajectories increase. Differentiating with respect to $h$ the Riccati equation in (5.18), we determine the curve of inflection points $R^{\prime \prime}(h)=0$ in the phase portrait, which is $(R-h)\left(R^{2}+2 h R-3\right)=0$. As $R^{\prime}=\frac{1}{4}$ on the line $R=h$, by (ii) and (iii) $\Gamma_{u}$ and $\Gamma_{l}$ cannot intersect this line. Hence, $\Gamma_{u}$ lies in the domain corresponding to the concave trajectories, and $\Gamma_{l}$ to the convex ones.

The proof of the statements listed below is obvious.
Lemma 7. The conic curve $w(h, R)=0$ has the following properties:
(i) It goes through the critical points $(1,1)$ and $(-1,-1)$ of $(5.18)$ and $w(h, h)=$ $6 \omega^{2}\left(h^{2}-1\right)$.
(ii) For $\omega=0$ it degenerates into $v^{2}(R-h)(4 R-h+3)=0$. For $\omega^{2}+2 \omega v-2 v^{2}=0$ it degenerates into $v^{2} h[(R-1+(7 \mp 4 \sqrt{3})(h-1)]=0$.
(iii) If $(h, R(h))$ are the local coordinates near $(1,1)$, then $R^{\prime}(1)=1-\frac{2 \omega^{2}}{v^{2}}, R^{\prime \prime}(1)=$ $\frac{8 \omega^{4}\left(\omega^{2}+2 \omega v-2 v^{2}\right)}{3 v^{6}}$.
(iv) If $(h(R), R)$ are the local coordinates near $(-1,-1)$ and $\omega \neq 0$, then $h^{\prime}(-1)=$ $0, h^{\prime \prime}(-1)=\frac{2 v^{2}-2 \omega v-\omega^{2}}{3 \omega^{2}}$.

After the preparation done above, we proceed to prove statements (i) and (ii) of proposition 8. To prove (i), let us denote by $\Omega$ the open triangle in the $(h, R)$ plane having vertices at $(1,1),(-1,-1)$ and $\left(1,-\frac{1}{2}\right)$. By lemma $6, \Gamma_{l} \subset \Omega$. On its turn, $\Omega \subset W_{-}$. This holds because by lemma 7 one has $w(h, h)<0$ for $|h|<1, w(1,1)=w(-1,-1)=$ $0, R^{\prime}(1)<1, h^{\prime}(-1)=0$, and because $w\left(1,-\frac{1}{2}\right)=-\frac{9}{2} \omega(\omega+2 v)<0$ (recall that $2 \omega>v>0$ in the truncated pendulum case). Proposition 8(i) is established.

The proof of (ii) is more complicated. In this case, $2 \omega<v$. Assume first that $\omega^{2}+2 \omega v-2 v^{2}>0$. Then, by lemma 7(iii), (iv) and convexity, the conic curve is a hyperbola. One of its branches is contained in the half-plane $h \leqslant-1$, the other branch lies
above its tangent line at $(1,1)$, namely $R-1=R^{\prime}(1)(h-1)$. As $R^{\prime}(1)<0$ in the considered case, by lemmas 6 (i) and 7 (i) we conclude that $\Gamma_{u} \subset W_{-}$. The same proof goes in the degenerate case $\omega^{2}+2 \omega v-2 v^{2}=0$.

The case $\omega^{2}+2 \omega v-2 v^{2}<0$ is more delicate. If $\omega=0$, then lemmas 6(i) and 7(ii) imply that $\Gamma_{u} \subset W_{+}$. Below we will take $\omega \neq 0$. Consider the branch $R=r(h)$ of the conic curve going through both $(1,1)$ and $(-1,-1)$. Writing $w(h, R)=\delta_{0} R^{2}+\delta_{1} R+\delta_{2}, \delta_{0}>0$, one obtains

$$
\begin{equation*}
r(h)=\frac{-\delta_{1}+\sqrt{D}}{2 \delta_{0}}, \quad D=\delta_{1}^{2}-4 \delta_{0} \delta_{2}>0, \quad|h|<1 \tag{5.20}
\end{equation*}
$$

Below we will denote this branch by $C$. By (5.19) and lemma 7(iv), we have near $h=-1$

$$
R+1 \sim-\frac{8}{\log (h+1)} \text { on } \Gamma_{u}, \quad R+1 \sim \sqrt{\frac{6 \omega^{2}(h+1)}{2 v^{2}-2 \omega v-\omega^{2}}} \text { on } C .
$$

This yields that for $h$ close to $-1, \Gamma_{u}$ is placed above $C$. Similarly, by lemmas 6(iii) and 7(iii) one obtains that near $h=1, \Gamma_{u}$ lies above $C$ if $3 v^{2}-8 \omega^{2} \geqslant 0$ and below $C$ otherwise. Therefore, to finish the proof of proposition 7, we have to establish the following: (i) $\Gamma_{u}$ is entirely placed above $C$ if $3 v^{2}-8 \omega^{2} \geqslant 0$ and (ii) $\Gamma_{u}$ intersects the conic just once if $3 v^{2}-8 \omega^{2}<0$. In the first case, we would have $\Gamma_{u} \subset W_{+}$, while in the second one, the part of $\Gamma_{u}$ near $h=1$ would be in $W_{-}$and the remaining part, in $W_{+}$. Unfortunately, both $\Gamma_{u}$ and $C$ are concave and it is not so easy to determine the number of their intersections.

Below, we proceed to determine the number of contact points that $C$ has with the vector field (5.18). Because of the type of critical points, in our case the number of intersections is less than or equal to the number of contact points. As is well known, the equation of the contact points is given by
$\frac{\mathrm{d}}{\mathrm{d} s}(R-r(h))_{\mid R=r(h)}=\left[\dot{R}-\dot{h} r^{\prime}(h)\right]_{\mid R=r(h)}=1-2 h r(h)+r^{2}(h)-4\left(1-h^{2}\right) r^{\prime}(h)=0$.

We replace $r^{\prime}=-\left(\delta_{1}^{\prime} r+\delta_{2}^{\prime}\right) /\left(2 \delta_{0} r+\delta_{1}\right)$ in (5.21) and use once again the quadratic equation satisfied by $r$ to obtain
$\left[\delta_{1}+4 \delta_{2} h+\frac{\delta_{1} \delta_{2}}{\delta_{0}}+4 \delta_{2}^{\prime}\left(1-h^{2}\right)\right]+r\left[2 \delta_{0}+2 \delta_{1} h-2 \delta_{2}+\frac{\delta_{1}^{2}}{\delta_{0}}+4 \delta_{1}^{\prime}\left(1-h^{2}\right)\right]=0$.
Next, replacing $r$ from (5.20) and performing the needed calculations we get

$$
\left[12 \delta_{0}\left(2 \omega v-v^{2}\right)\left(1-h^{2}\right)+D\right] \sqrt{D}-\left[3 v^{2}(1+h) D+2 \delta_{0}\left(1-h^{2}\right) D^{\prime}\right]=0
$$

By (5.20) and (5.17) one has
$D=(1+h)^{2}\left(9 v^{4}+24 \omega^{2} \delta_{0} \zeta\right), \quad D^{\prime}=(1+h)\left(18 v^{4}+24 \omega^{2} \delta_{0}(\zeta-1)\right), \quad \zeta=\frac{1-h}{1+h}>0$.
We replace these values in the equation above and divide the result by $(1+h)^{3}$. One obtains

$$
\begin{aligned}
{\left[9 v^{4}+12\left(2 \omega^{2}\right.\right.} & \left.\left.+2 \omega v-v^{2}\right) \delta_{0} \zeta\right] \sqrt{9 v^{4}+24 \omega^{2} \delta_{0} \zeta} \\
& -\left[27 v^{6}+12\left(3 v^{4}+6 \omega^{2} v^{2}-4 \omega^{2} \delta_{0}\right) \delta_{o} \zeta+48 \omega^{2} \delta_{0}^{2} \zeta^{2}\right]=0
\end{aligned}
$$

At the end, we introduce a new variable $z$ by

$$
\sqrt{9 v^{4}+24 \omega^{2} \delta_{0} \zeta}=3\left(v^{2}+z\right), \quad z \in(0, \infty)
$$

to obtain the equation of contact points in a final form $P(z)=0$, where

$$
\begin{aligned}
P(z)=3 z^{3}+ & 6\left(3 v^{2}-2 \omega v-2 \omega^{2}\right) z^{2} \\
& +4(v+\omega)\left(9 v^{3}-18 \omega v^{2}+4 \omega^{2} v+4 \omega^{3}\right) z+2 v^{2} \delta_{0}\left(3 v^{2}-8 \omega^{2}\right)
\end{aligned}
$$

Taking $P(z)=3 z^{3}+A_{2} z^{2}+A_{1} z+A_{0}$, it is easy to verify that

$$
\begin{array}{ll}
A_{2}>0, A_{1}>0, A_{0}>0 & \text { for } \quad 3 v^{2}-8 \omega^{2}>0 \\
A_{2}>0, A_{1}>0, A_{0}<0 & \text { for } \quad 3 v^{2}-8 \omega^{2}<0, v+\omega>0 \\
A_{1}<0, A_{0}<0 & \text { for } \quad v+\omega<0
\end{array}
$$

(recall that $v-2 \omega>0$ and $\delta_{0}=4 v^{2}-4 \omega v-2 \omega^{2}>0$ ). By Descartes chain rule, the equation $P(z)=0$ has no positive root if $3 v^{2}-8 \omega^{2} \geqslant 0$ and has exactly one positive root if $3 v^{2}-8 \omega^{2}<0$ and $2 v^{2}-2 \omega v-\omega^{2}>0$. All changes of the variables we used throughout the proof were one-to-one, hence there is no contact point in the first case and there is just one contact point in the second case. As a result, in the first case $C$ does not intersect $\Gamma_{u}$ and therefore $\Gamma_{u} \subset W_{+}$. In the second case, $C$ intersects $\Gamma_{u}$ at a unique point corresponding to some $\bar{h}_{0} \in(-1,1)$ so that the part of $\Gamma_{u}$ related to $\left(-1, \bar{h}_{0}\right)$ is in $W_{+}$and the remaining part is in $W_{-}$. Let $\varphi_{0}$ be the value corresponding to $h_{0}$ according to (3.11) and let $T_{\max }$ be the period of the orbit going through $\left(\varphi_{0}, 0\right)$ in the $\left(\varphi, \varphi^{\prime}\right)$ plane. Then the orbits from the period annulus having a period $T>T_{\max }$ will be stable and the remaining ones unstable. Proposition 8 is completely proved.

Let us recall again that in the left(right) Duffing oscillator case the period $T$ belongs to the interval $I=\left(T_{0}, \infty\right)$ where $T_{0}=2 \pi \sqrt{\frac{v}{2 v-4 \omega}}$. Lemma 3 and proposition 8 imply the following.

Corollary 2. Both conditions of assumption 1 are satisfied in the left (right) Duffing oscillator case, provided that
(i) $3 v^{2}-8 \omega^{2} \geqslant 0$;
(ii) $3 v^{2}-8 \omega^{2}<0,2 v^{2}-2 \omega v-\omega^{2}>0$ and the period $T$ is sufficiently large.

Proof of theorem I. Theorem I is a direct consequence of corollary 2 and theorem 2, taking into account that $\varphi$ does not oscillate around zero only in the left and right Duffing oscillator cases.

Example 5.3 (coherent single-power nonlinearities). Let $a(u)=\beta u^{k+1}, b(u)=\gamma \beta u^{k}, k \in \mathbb{N}$, where $|\gamma|<1$ and $v / \beta>0$. Or more generally, one can take $a(u)=2 \omega u+\beta u^{k+1}, b(u)=$ $2 \omega+\gamma \beta u^{k}$, where $r=(v-2 \omega) / \beta>0$. For simplicity, assume that $c_{1}=0$. Then (3.5) is satisfied with $X_{0}=r^{1 / k}$ and proposition 1 works. Next, equation (3.3) implies that $\varphi$ depends on $r$ alone. Moreover, one has
$\varphi=r^{1 / k} \bar{\varphi}, \quad \varphi_{0}=r^{1 / k} \bar{\varphi}_{0}, \quad \varphi_{1}=r^{1 / k} \bar{\varphi}_{1}, \quad c_{2}=r^{\frac{k+2}{k}} \bar{c}_{2}, \quad U(\varphi)=r^{\frac{2}{k}} \bar{U}(\bar{\varphi})$,
where $\bar{\varphi}, \bar{\varphi}_{0}, \bar{\varphi}_{1}, \bar{c}_{2}$ and $\bar{U}$ do not depend on $v, \omega$ and $\beta$. Therefore, we obtain

$$
\begin{align*}
\ddot{\mathrm{d}}(v) & =\frac{1}{2} \frac{\mathrm{~d}}{\mathrm{~d} v} \int_{0}^{T}\left(\varphi^{2}+\varphi^{\prime 2}\right) \mathrm{d} x=\left(\frac{1}{2} \frac{\mathrm{~d}}{\mathrm{~d} v} r^{2 / k}\right) \int_{0}^{T}\left(\bar{\varphi}^{2}+\bar{\varphi}^{\prime 2}\right) \mathrm{d} x \\
& =\frac{r^{2 / k}}{k(v-2 \omega)} \int_{0}^{T}\left(\bar{\varphi}^{2}+\bar{\varphi}^{\prime 2}\right) \mathrm{d} x=\frac{1}{k(v-2 \omega)} \int_{0}^{T}\left(\varphi^{2}+\varphi^{\prime 2}\right) \mathrm{d} x . \tag{5.22}
\end{align*}
$$

Therefore, $\ddot{\mathrm{d}}(v)$ takes the sign of $\beta$ (and $v-2 \omega$ ).

## 6. The perturbed gBBM equation

Let us denote by $\mathcal{C}(X, Y)$ (respectively, $\mathcal{B}(X, Y)$ ) the set of all closed (respectively, bounded) operators $S: X \rightarrow Y$. When $X=Y$, we shall write simply $\mathcal{C}(X)$ and $\mathcal{B}(X)$. If $S \in \mathcal{C}(X, Y)$, denote by $\mathbf{G}(S) \subset X \times Y$ its graph. Below we choose $\left.X=Y=L^{2}[0, T]\right)$. One can define a metric $\hat{\delta}$ on $\mathcal{C}\left(L^{2}[0, T]\right)$ as follows: for $R, S \in \mathcal{C}\left(L^{2}[0, T]\right)$, set

$$
\hat{\delta}(R, S)=\left\|P_{R}-P_{S}\right\|_{\mathcal{B}\left(L^{2} \times L^{2}\right)}
$$

where $P_{R}$ and $P_{S}$ are the orthogonal projections on the graphs of $\mathbf{G}(R)$ and $\mathbf{G}(S)$, respectively. The following results are well known (see Kato [39, chapter 4, theorems 2.14 and 2.17]):

Theorem A. Take $X=Y=L^{2}[0, T]$ and assume that $R, S \in \mathcal{C}(X, Y), A \in \mathcal{B}(X, Y)$. Then

$$
\hat{\delta}(R+A, S+A) \leqslant 2\left(1+\|A\|^{2}\right) \hat{\delta}(R, S)
$$

Theorem B. Assume that $S \in \mathcal{C}(X, Y)$ and $B$ is a $S$-bounded operator satisfying $\|B f\| \leqslant$ $a\|f\|+b\|S f\|$ with $b<1$. Then

$$
R=S+B \in \mathcal{C}(X, Y) \quad \text { and } \quad \hat{\delta}(R, S) \leqslant \frac{\sqrt{a^{2}+b^{2}}}{1-b}
$$

Let us introduce a small perturbation in (gBBM) by taking $a(u)=2 \omega u+\beta u^{k}, k=2,3$ and $b(u)=\gamma g(u)$ in the general equation (1.1) where $\gamma$ is a small real parameter. We will denote the travelling-wave solution corresponding to (1.1) again by $\varphi_{v}$ and the one corresponding to (gBBM) (that is, when $\gamma=0$ ) by $\varphi_{v}^{0}$. If $T$ is the period of $\varphi_{v}^{0}$, then by proposition 3 for sufficiently small $\gamma$ there is a smooth function $\varphi_{0}(\gamma)$ such that the wave solution $\varphi_{v}$ of the perturbed equation which goes through the point $\left(\varphi_{0}(\gamma), 0\right)$ in the $\left(\varphi, \varphi^{\prime}\right)$ plane will have the same period $T$. For the related operators $\mathcal{H}_{v}$, we have respectively
$\mathcal{H}_{v}=-v \partial_{x}^{2}+v-a^{\prime}\left(\varphi_{v}\right)+\gamma\left[g\left(\varphi_{v}\right) \partial_{x}^{2}+g^{\prime}\left(\varphi_{v}\right) \varphi_{v}^{\prime} \partial_{x}+\frac{1}{2} g^{\prime \prime}\left(\varphi_{v}\right) \varphi_{v}^{\prime 2}+g^{\prime}\left(\varphi_{v}\right) \varphi_{v}^{\prime \prime}\right]$,
$\mathcal{H}_{v}^{0}=-v \partial_{x}^{2}+v-a^{\prime}\left(\varphi_{v}^{0}\right)$.
Theorem 3. If assumption 1(i) holds for the operator $\mathcal{H}_{v}^{0}$, then for $\gamma$ sufficiently small it also holds for the operator $\mathcal{H}_{v}$.

Proof. In order to apply theorem A, we define the operator $A$ to be a multiplication by the function $-a^{\prime}\left(\varphi_{v}^{0}\right)$ and take $S=-v \partial_{x}^{2}+v$. Then denoting $B_{\gamma}=\mathcal{H}_{v}-\mathcal{H}_{v}^{0}$, we obtain from theorem A the estimate

$$
\begin{equation*}
\hat{\delta}\left(\mathcal{H}_{v}, \mathcal{H}_{v}^{0}\right)=\hat{\delta}\left(S+B_{\gamma}+A, S+A\right) \leqslant 2\left(1+\|A\|^{2}\right) \hat{\delta}\left(S+B_{\gamma}, S\right) . \tag{6.1}
\end{equation*}
$$

As

$$
B_{\gamma}=a^{\prime}\left(\varphi_{v}^{0}\right)-a^{\prime}\left(\varphi_{v}\right)+\gamma\left[-v^{-1} g\left(\varphi_{v}\right) S+g^{\prime}\left(\varphi_{v}\right) \varphi_{v}^{\prime} \partial_{x}+G\right]
$$

where

$$
G=g\left(\varphi_{v}\right)+\frac{1}{2} g^{\prime \prime}\left(\varphi_{v}\right) \varphi_{v}^{\prime 2}+g^{\prime}\left(\varphi_{v}\right) \varphi_{v}^{\prime \prime}
$$

for $f \in D(S)=H^{2}([0, T])$, we further have

$$
\begin{aligned}
\left\|B_{\gamma} f\right\| \leqslant \max _{[0, T]} & \left|a^{\prime}\left(\varphi_{v}^{0}\right)-a^{\prime}\left(\varphi_{v}\right)\right| \cdot\|f\| \\
& +|\gamma|\left[\max _{[0, T]}\left|g\left(\varphi_{v}\right) / v\right| \cdot\|S f\|+\max _{[0, T]}\left|g^{\prime}\left(\varphi_{v}\right) \varphi_{v}^{\prime}\right| \cdot\left\|\partial_{x} f\right\|+\max _{[0, T]}|G| \cdot\|f\|\right] .
\end{aligned}
$$

From Plancherel identity, we have

$$
\begin{aligned}
\|S f\|^{2} & =\int_{0}^{T} S f \cdot \overline{S f} \mathrm{~d} x=T \sum_{k=-\infty}^{+\infty}|\widehat{S f}(k)|^{2} \\
& =v^{2} \sum_{k=-\infty}^{+\infty}\left(1+|k|^{2}\right)^{2}|\bar{f}(k)|^{2} \geqslant K(v) \sum_{k=-\infty}^{+\infty}|k|^{2}|\bar{f}(k)|^{2}=K(v)\left\|\partial_{x} f\right\|^{2}
\end{aligned}
$$

where $K(v)$ is a constant depending only on $v$. By using propositions 3 and 4, we obtain that $\max _{[0, T]}\left|a^{\prime}\left(\varphi_{v}^{0}\right)-a^{\prime}\left(\varphi_{v}\right)\right|=O(\gamma)$. Taking into account these estimates, we come to the conclusion that the operator $B=B_{\gamma}$ is $S$-bounded, with constants $a, b$ tending to zero as $\gamma \rightarrow 0$. Applying theorem B and inequality (6.1), we obtain

$$
\begin{equation*}
\hat{\delta}\left(\mathcal{H}_{v}, \mathcal{H}_{v}^{0}\right) \leqslant 2\left(1+\|A\|^{2}\right) \frac{\sqrt{a^{2}+b^{2}}}{1-b}=O(\gamma) . \tag{6.2}
\end{equation*}
$$

Therefore, by [39] (theorem 3.16 in chapter 4), the operators $\mathcal{H}_{v}$ and $\mathcal{H}_{v}^{0}$ have the same spectral properties. More precisely (see also [1, p 363]), let $U$ be the open disc bounded by $\Gamma$ and $\operatorname{spec}\left(\mathcal{H}_{v}^{0}\right) \cap \bar{U}=\left\{\alpha_{0}, 0\right\}$, where $\alpha_{0}$ is the negative eigenvalue of $\mathcal{H}_{v}^{0}$. Choose circular contours $\Gamma_{1}$ and $\Gamma_{2}$ contained in $U$, such that if $U_{1}$ and $U_{2}$ are the open discs bounded by $\Gamma_{1}$ and $\Gamma_{2}$, respectively, then $\operatorname{spec}\left(\mathcal{H}_{v}^{0}\right) \cap \bar{U}_{1}=\left\{\alpha_{0}\right\}$ and $\operatorname{spec}\left(\mathcal{H}_{v}^{0}\right) \cap \bar{U}_{2}=\{0\}$. From (6.2) for sufficiently small $\gamma$ we have that the $\operatorname{spec}\left(\mathcal{H}_{v}\right) \cap \bar{U}_{1}$ and $\operatorname{spec}\left(\mathcal{H}_{v}\right) \cap \bar{U}_{2}$ consist of a single, simple eigenvalue. Since 0 is an eigenvalue of $\mathcal{H}_{v}$ we must have $\operatorname{spec}\left(\mathcal{H}_{v}\right) \cap \bar{U}_{2}=\{0\}$, which shows that 0 is a simple eigenvalue of $\mathcal{H}_{v}$. Similarly we obtain that $\operatorname{spec}\left(\mathcal{H}_{v}\right) \cap \bar{U}$ consists of a finite set of eigenvalues of total multiplicity 2 . This completes the proof of the theorem.

Proof of theorems II and III. Take $\omega=0$. Equality (5.22) implies that $\ddot{\mathrm{d}}(v)$ is positive. When $\gamma=0$, applying respectively proposition 7 and corollary 1 , we conclude that $\mathcal{H}_{v}^{0}$ has the needed spectral properties in both cases. By theorem 3 the same remains true for small $|\gamma|$ as well. Hence, theorem 2 applies to both cases.

## 7. Small-amplitude travelling-wave solutions

Since the expressions in (3.10) and (4.9) are too complex, it is difficult to use them in general. Below we are going to consider the simpler case when the periodic waves we study are of small amplitude. That is, $\varphi_{1}-\varphi_{0}$ is close to zero and therefore the periodic trajectory is entirely contained in a small neighbourhood of some centre $\left(X_{0}, 0\right) \in \mathbb{R}^{2}$ given by (3.5). Below we establish that an important role in the stability of the small-amplitude travelling-wave solutions is played by the first isochronous (or period) constant. Let us recall that the period function has an expansion

$$
T(r)=T_{0}+T_{2} r^{2}+T_{4} r^{4}+T_{6} r^{6}+T_{8} r^{8}+\cdots
$$

with respect to $r$, the distance between the centre at $\left(X_{0}, 0\right)$ and the intersection point of the orbit with the $x$-axis $\left(\varphi_{0}, 0\right)$. Then $T_{2 k}, k=1,2, \ldots$, is the $k$ th isochronous constant. When all period constants are zero, the centre is isochronous and all orbits surrounding it have the same period $T_{0}$. Since $T^{\prime}\left(\varphi_{0}\right)=-\mathrm{d} T / \mathrm{d} r$, the above expansion yields immediately the following proposition.

Proposition 9. The period of a small-amplitude travelling-wave solution around a nonisochronous centre is not critical.

Let us mention, however, that the calculation of the isochronous constants is a difficult task and the problem of isochronicity is completely solved for only few particular cases. Denote for short
$p=-\frac{a^{\prime \prime}\left(X_{0}\right)}{3\left(a^{\prime}\left(X_{0}\right)-v\right)}, \quad q=-\frac{a^{\prime \prime \prime}\left(X_{0}\right)}{3\left(a^{\prime}\left(X_{0}\right)-v\right)}, \quad m=\frac{b^{\prime}\left(X_{0}\right)}{v-b\left(X_{0}\right)}, \quad n=\frac{b^{\prime \prime}\left(X_{0}\right)}{v-b\left(X_{0}\right)}$.

As we shall see below, the first isochronous constant in our case is expressed by
$T_{2}=T_{0} \vartheta, \quad T_{0}=2 \pi \sqrt{\frac{v-b\left(X_{0}\right)}{a^{\prime}\left(X_{0}\right)-v}}, \quad \vartheta=\frac{1}{16}\left(15 p^{2}+3 q-m^{2}-6 p m-2 n\right)$.
Take a small positive $\varepsilon$ and let $\varphi_{0}=X_{0}-\varepsilon$. Then by (3.9) we obtain

$$
\begin{equation*}
\varphi_{1}=X_{0}+\varepsilon+p \varepsilon^{2}+p^{2} \varepsilon^{3}+O\left(\varepsilon^{4}\right) \tag{7.2}
\end{equation*}
$$

Expression (7.2) as well as the formulae that follow are obtained by long and boring asymptotic calculations which we will omit here. Thus, by using (3.8) we come to the expression
$U(X)=\frac{\varepsilon^{2}\left(a^{\prime}\left(X_{0}\right)-v\right)\left(1-s^{2}\right)\left[1+p(1-s) \varepsilon-\frac{1}{4}\left(p^{2}(1+6 s)+q\left(1+s^{2}\right)\right) \varepsilon^{2}+\cdots\right]}{\left(v-b\left(X_{0}\right)\right)\left[1-m s \varepsilon-\frac{1}{2}\left(m p(1+s)+n s^{2}\right) \varepsilon^{2}+\cdots\right]}$.
Then, by (3.7) and (3.8) we obtain

$$
\begin{equation*}
T=T_{0}\left[1+\vartheta \varepsilon^{2}+O\left(\varepsilon^{4}\right)\right] . \tag{7.3}
\end{equation*}
$$

In order to calculate the sign of the derivative $\ddot{\mathrm{d}}(v)$, we prefer to apply a more direct approach, deriving first an alternative formula instead of (4.9). By using (4.3), we obtain

$$
\ddot{\mathrm{d}}(v)=\frac{\mathrm{d}}{\mathrm{~d} v}\left(\int_{0}^{T / 2}\left(\varphi^{2}+\varphi^{\prime 2}\right) \mathrm{d} x\right)=\frac{\mathrm{d}}{\mathrm{~d} v}\left(\int_{\varphi_{0}}^{\varphi_{1}}\left[s^{2}+U(s)\right] \frac{\mathrm{d} s}{\sqrt{U(s)}}\right) .
$$

In the calculation below, we denote by $\dot{\varphi}_{0}, \dot{\varphi}_{1}$, etc, the derivatives with respect to the parameter $v$. As $U(s)$ vanishes at $\varphi_{0}$ and $\varphi_{1}$, then clearly

$$
\frac{\mathrm{d}}{\mathrm{~d} v}\left(\int_{\varphi_{0}}^{\varphi_{1}} \sqrt{U(s)} \mathrm{d} s\right)=\int_{\varphi_{0}}^{\varphi_{1}} \frac{\left(\dot{U}(s)+\dot{\varphi}_{0} U_{\varphi_{0}}(s)\right) \mathrm{d} s}{2 \sqrt{U(s)}} .
$$

To perform differentiation in the other part of the integral, we first use a change of the variable like (3.8) and return to the initial variable after the calculations are made. Thus,

$$
\begin{aligned}
& \frac{\mathrm{d}}{\mathrm{~d} v}\left(\int_{\varphi_{0}}^{\varphi_{1}} \frac{s^{2} \mathrm{~d} s}{\sqrt{U(s)}}\right)=\frac{\dot{\varphi}_{1}-\dot{\varphi}_{0}}{\varphi_{1}-\varphi_{0}} \int_{\varphi_{0}}^{\varphi_{1}} \frac{s^{2} \mathrm{~d} s}{\sqrt{U(s)}} \\
& +\int_{\varphi_{0}}^{\varphi_{1}} \frac{\left(\dot{\varphi}_{1}-\dot{\varphi}_{0}\right) s+\dot{\varphi}_{0} \varphi_{1}-\dot{\varphi}_{1} \varphi_{0}}{\varphi_{1}-\varphi_{0}} \frac{2 s \mathrm{~d} s}{\sqrt{U(s)}} \\
& -\int_{\varphi_{0}}^{\varphi_{1}}\left(\dot{U}(s)+\dot{\varphi}_{0} U_{\varphi_{0}}(s)+\frac{\left(\dot{\varphi}_{1}-\dot{\varphi}_{0}\right) s+\dot{\varphi}_{0} \varphi_{1}-\dot{\varphi}_{1} \varphi_{0}}{\varphi_{1}-\varphi_{0}} U^{\prime}(s)\right) \frac{s^{2} \mathrm{~d} s}{2 U^{3 / 2}(s)} .
\end{aligned}
$$

As a result, one obtains

$$
\begin{equation*}
\ddot{\mathrm{d}}(v)=\int_{\varphi_{0}}^{\varphi_{1}} \frac{V(s) \mathrm{d} s}{2 U^{3 / 2}(s)}, \tag{7.4}
\end{equation*}
$$

with

$$
\begin{align*}
& V(s)=2 \frac{\dot{\varphi}_{1}-\dot{\varphi}_{0}}{\varphi_{1}-\varphi_{0}} s^{2} U(s)+\left(U(s)-s^{2}\right)\left(\dot{U}(s)+\dot{\varphi}_{0} U_{\varphi_{0}}(s)\right) \\
&+\left(4 s U(s)-s^{2} U^{\prime}(s)\right) \frac{\left(\dot{\varphi}_{1}-\dot{\varphi}_{0}\right) s+\dot{\varphi}_{0} \varphi_{1}-\dot{\varphi}_{1} \varphi_{0}}{\varphi_{1}-\varphi_{0}} \tag{7.5}
\end{align*}
$$

From (3.5), (7.1) and (7.2), for the derivatives with respect to $v$, we get respectively

$$
\begin{align*}
\dot{\varphi}_{0} & =\frac{X_{0}}{a^{\prime}\left(X_{0}\right)-v}-\dot{\varepsilon} \\
\dot{\varphi}_{1} & =\frac{X_{0}+\left(p+3 p^{2} X_{0}+q X_{0}\right) \varepsilon^{2}+O\left(\varepsilon^{3}\right)}{a^{\prime}\left(X_{0}\right)-v}+\left[1+2 p \varepsilon+3 p^{2} \varepsilon^{2}+O\left(\varepsilon^{3}\right)\right] \dot{\varepsilon} \tag{7.6}
\end{align*}
$$

Replacing (7.2) and (7.6) in the expression of $V(X)$ given by (7.5), one obtains

$$
\begin{aligned}
V(X)=\frac{\left(1-s^{2}\right) X_{0}^{2}}{v-}\left(5\left(X_{0}\right)\right. & \left.5+(3 p-m) X_{0}+\frac{a^{\prime}\left(X_{0}\right)-v}{v-b\left(X_{0}\right)}\right) \varepsilon^{2}+O\left(\varepsilon^{3}\right) \\
& +\frac{\left(1-s^{2}\right)\left(a^{\prime}\left(X_{0}\right)-v\right) \varepsilon^{2} \dot{\varepsilon}}{v-b\left(X_{0}\right)}\left[\left(p X_{0}^{2}+4 X_{0}-m X_{0}^{2}\right) s\right. \\
& +\left(4 s^{2}+\frac{2\left(a^{\prime}\left(X_{0}\right)-v\right)}{v-b\left(X_{0}\right)}+X_{0}\left[p\left(4+8 s-2 s^{2}\right)+2 m s^{2}\right]+X_{0}^{2}\left[p^{2}(3+2 s)\right.\right. \\
& \left.\left.\left.+p m\left(2 s^{2}-2 s-1\right)+\frac{q}{2}\left(1+s^{2}\right)-\left(2 m^{2}+n\right) s^{2}\right]\right) \varepsilon+O\left(\varepsilon^{2}\right)\right] .
\end{aligned}
$$

The leading term of $\dot{\varepsilon}$ can be determined from (7.3) after differentiation with respect to $v$. One obtains

$$
\dot{\varepsilon} \sim-\frac{\dot{T}_{0}}{2 T_{2} \varepsilon}=\frac{1}{4\left(a^{\prime}\left(X_{0}\right)-v\right) \vartheta \varepsilon}\left(X_{0} m-3 p X_{0}-1-\frac{a^{\prime}\left(X_{0}\right)-v}{v-b\left(X_{0}\right)}\right)
$$

Finally, replacing in integral (7.4), we derive the formula

$$
\begin{equation*}
\ddot{\mathrm{d}}(v)=\left(a^{\prime}\left(X_{0}\right)-v\right)\left[\frac{T_{0} X_{0}^{2}}{\left(a^{\prime}\left(X_{0}\right)-v\right)^{2}}-\frac{\dot{T}_{0}^{2}}{2 T_{2}}\right]+O(\varepsilon) \tag{7.7}
\end{equation*}
$$

Example 1 (the BBM equation). Let $a(u)=2 \omega u+\frac{3}{2} u^{2}, b(u)=0$. Assume for definiteness that $v>0$. The other case is considered similarly. Using the notation introduced earlier and the expression of $X_{0}$, we obtain
$a^{\prime}\left(X_{0}\right)-v=D^{1 / 2}, \quad p=-\frac{1}{D^{1 / 2}}, \quad q=m=n=0, \quad \vartheta=\frac{15}{16 D}, \quad T_{0}=2 \pi \frac{v^{1 / 2}}{D^{1 / 4}}$,
$T_{2}=\frac{15 T_{0}}{16 D}, \quad \dot{T}_{0}=\frac{T_{0}\left(2 \omega^{2}-\omega v+3 c_{1}\right)}{v D}, \quad \ddot{\mathrm{~d}}(v)=\frac{T_{0}}{D^{1 / 2}}\left[X_{0}^{2}-\frac{8\left(2 \omega^{2}-\omega v+3 c_{1}\right)^{2}}{15 v^{2}}\right]$.
Now, expressing $c_{1}$ and $X_{0}$ through $D$, we rewrite the expression in the brackets of $\ddot{\mathrm{d}}(v)$ as

$$
\begin{aligned}
\Lambda= & X_{0}^{2}-\frac{2\left(D-v^{2}+2 \omega v\right)^{2}}{15 v^{2}} \\
= & -\frac{2}{15 v^{2}}\left[D+\sqrt{\frac{5}{6}} v D^{1 / 2}-\left(1-\sqrt{\frac{5}{6}}\right) v(v-2 \omega)\right] \\
& \times\left[D-\sqrt{\frac{5}{6}} v D^{1 / 2}-\left(1+\sqrt{\frac{5}{6}}\right) v(v-2 \omega)\right]
\end{aligned}
$$

Next, denote by $\lambda_{1}, \lambda_{2}$ the roots (with respect to $D^{1 / 2}$ ) of the first multiplier, and respectively by $\lambda_{3}, \lambda_{4}$ the roots of the second multiplier.
(a) If $0<v \leqslant \frac{48}{361}\left(9+5 \sqrt{\frac{5}{6}}\right) \omega$, then both multipliers are positive and $\Lambda<0$.
(b) If $\frac{48}{361}\left(9+5 \sqrt{\frac{5}{6}}\right) \omega<v \leqslant 2 \omega$, then the first multiplier is positive, $0 \leqslant \lambda_{4}<\lambda_{3}$ and therefore $\Lambda>0 \Leftrightarrow \lambda_{4}<D^{1 / 2}<\lambda_{3}$.
(c) If $v>2 \omega$, then $0<\lambda_{1}<\lambda_{3}, \lambda_{2}<0, \lambda_{4}<0$. Therefore $\Lambda>0 \Leftrightarrow \lambda_{1}<D^{1 / 2}<\lambda_{3}$.

Taking a second power of the above inequalities, it follows that

As $\lambda_{k}$ are homogeneous of first degree with respect to $v, \omega$, one can reformulate (7.8) as $\ddot{\mathrm{d}}(v)>0$ if and only if $\left(\omega / v, c_{1} / v^{2}\right) \in \Omega \subset \mathbb{R}^{2}$ where $\Omega$ can be explicitly written down if needed by using the formulae of the quadratic roots $\lambda_{k}$. We are not going to do this.

Example 2 (the modified BBM equation). Let $a(u)=2 \omega u+\beta u^{3}, b(u)=0, c_{1}=0$.
(i) $X_{0}=0$ (global centre or truncated pendulum case). Then $\ddot{\mathrm{d}}(v)=\frac{4 \omega^{2}}{3 \beta v^{2}} T_{0}+O(\varepsilon)$ and $\vartheta=\frac{3 \beta}{8(v-2 \omega)} \neq 0$.
(ii) $X_{0}^{2}=\frac{v-2 \omega}{\beta}$ (Duffing oscillator). Then $\ddot{\mathrm{d}}(v)=\frac{3 v^{2}-8 \omega^{2}}{6 \beta v^{2}} T_{0}+O(\varepsilon)$ and $\vartheta=\frac{3 \beta}{4(v-2 \omega)}>0$.

As seen from these examples, $\mathrm{d}(v)$ could be negative or positive, depending on the case.
Proof of theorems IV and $V$. In examples 1 and 2 , we calculated the first term of $\ddot{\mathrm{d}}(v)$ provided that $b(u)=0$ and understood when it is positive. If one take $b(u)=\gamma g(u)$ with $|\gamma|$ small, then $\ddot{\mathrm{d}}(v)$ will differ from the case $b(u)=0$ by a term which is $O(\gamma)$. Therefore, $\ddot{\mathrm{d}}(v)$ will keep the sign of its first term provided that $|\gamma|$ and the amplitude $\varepsilon$ are small enough. This means that $\ddot{\mathrm{d}}(v)>0$ in the domain $\Omega$ from the first example and for $3 v^{2}-8 \omega^{2}>0$ in the Duffing oscillator case of the second example (recall that all solutions not oscillating around zero belong to this case). Since the operator $\mathcal{H}_{v}^{0}$ related to the above cases satisfies the spectral properties required in assumption 1, the statements of theorems IV and V follow from theorems 2 and 3.

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